

Keynote talk: Modeling and Bias



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Have you seen this ILRS webpage?

Contributions to Earth Science - Windows Internet Explorer

ILRS http://ilrs.gsfc.nasa.gov/science_analysis/contributions.html Google


お気に入り ILRS Contributions to Earth Science

Analysis Centers

- Analysis Products
- Analysis Reports
- ILRS Analysis Center Conventions
- ITRF Solutions
- IERS Conventions Center
- IERS Conventions Update
- Contributions to Earth Science
- ILRS Bibliography
- ILRS Papers
- SLR and Earth Science
- Science Meetings

Satellite Laser Ranging Contributions to Earth Science

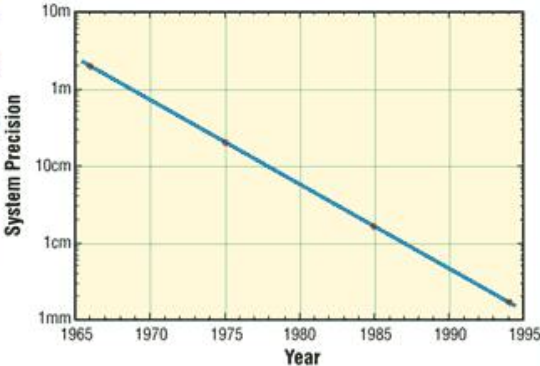
In Satellite Laser Ranging, a short pulse of coherent light generated by a laser (Light Amplification by Stimulated Emission of Radiation) is transmitted in a narrow beam to illuminate corner cube retroreflectors on the satellite. The return signal, typically a few photons, is collected by a telescope and the time-of-flight is measured. Using information about the satellite's orbit, the time-of-flight, and the speed of light, the location of the ranging station can be determined. Similar data acquired by another station, many kilometers distant from the first, or on a different continent, can be used to determine the distance between stations to precisions of centimeters or better. Repetitive measurements over months and years yield the change in distance, or the motion of the Earth's crust.



Laser ranging to a near-Earth satellite was initiated by NASA in 1964 with the launch of the Beacon-B satellite. Since that time, ranging precision, spurred by scientific requirements, has improved by a factor of a thousand from a few meters to a few millimeters. Similarly, the network of laser stations has grown from a few experimental sites to a global network of 43 stations in more than 30 countries. Most of these stations (33) are funded by organizations other than NASA.

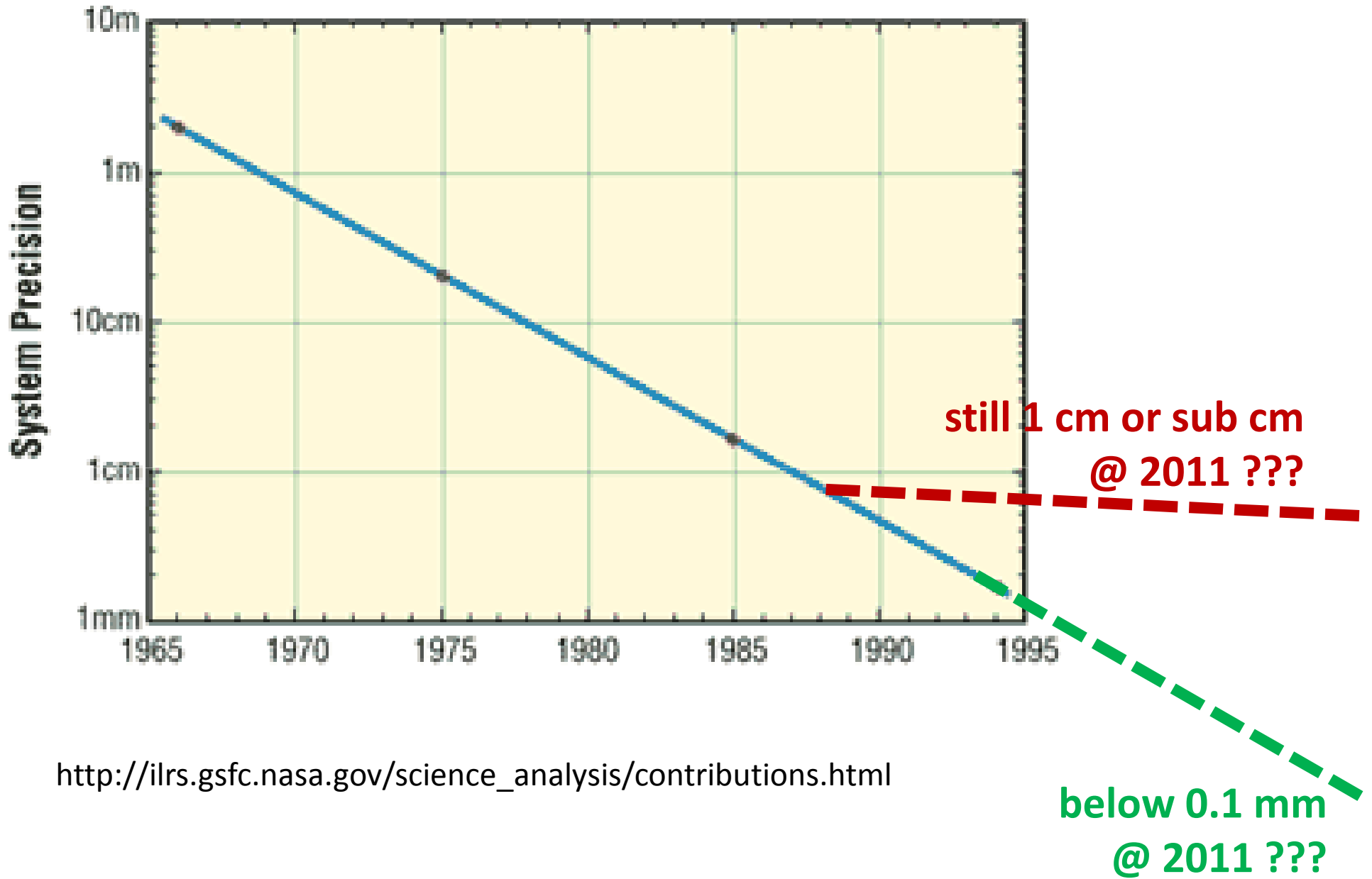
During the past three decades, the global Satellite Laser Ranging (SLR) network has evolved into a powerful source of data for studies of the solid Earth and its ocean and atmospheric systems including:

- Detection and monitoring of tectonic plate motion, crustal deformation, Earth rotation, and polar motion;



Year	System Precision (m)
1965	2
1970	1
1975	0.5
1980	0.25
1985	0.125
1990	0.0625
1995	0.03125

http://ilrs.gsfc.nasa.gov/science_analysis/contributions.html



http://ilrs.gsfc.nasa.gov/science_analysis/contributions.html

1 mm, sub-mm achievable?

Random error (precision)

kHz ranging has reduced the NP precision

Single-shot 5 mm RMS / $\text{Sqrt}[10000 \text{ (shots/NP)}] = 0.05 \text{ mm (NP)}!!$

Systematic/offset error (accuracy)

timer, laser, calibration, local survey, etc.

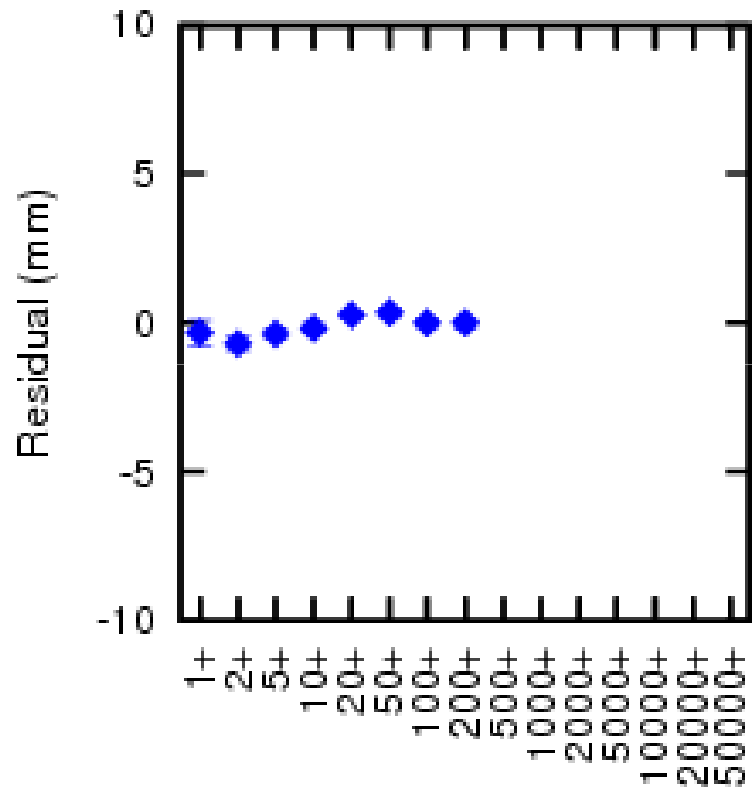
atmospheric delay

- ➔ intensity dependence
- ➔ microwave vs laser
- ➔ centre-of-mass correction
- ➔ modeling in orbit analysis

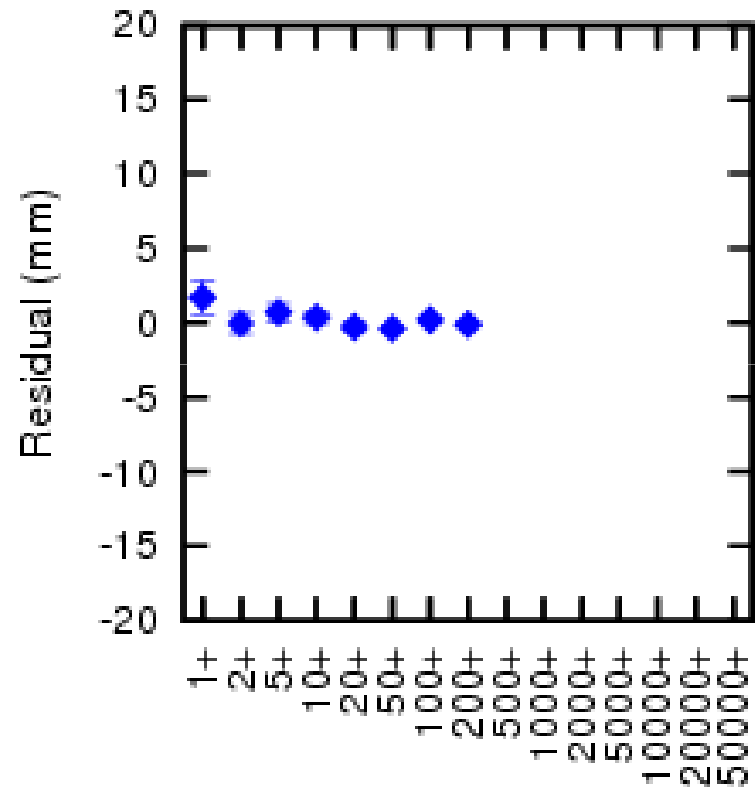
➔ **all affecting the quality of analysis products**

Bias vs Intensity 2010

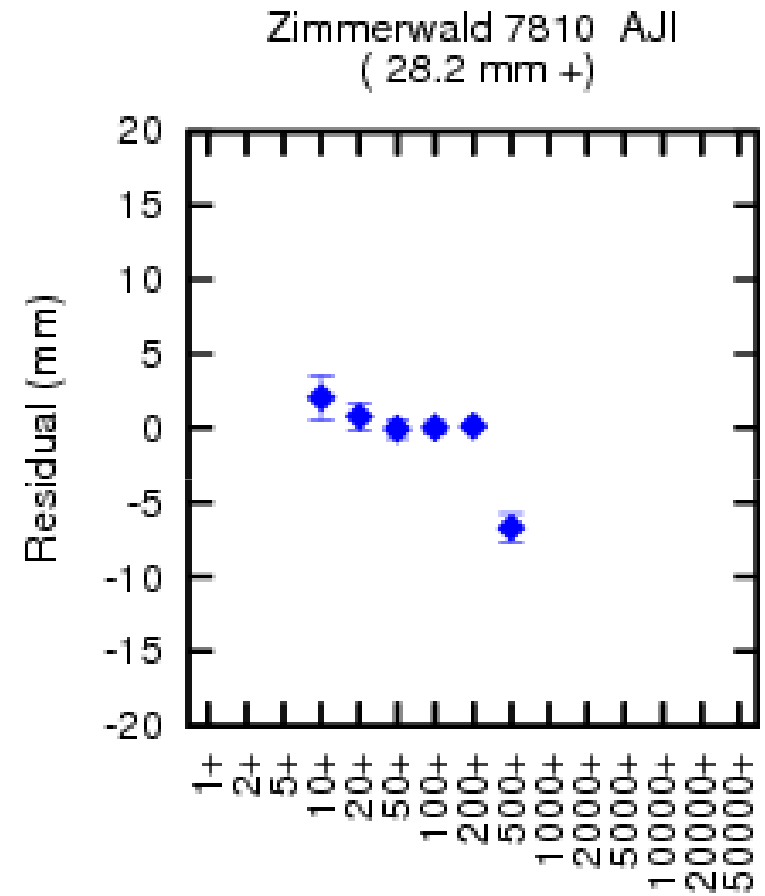
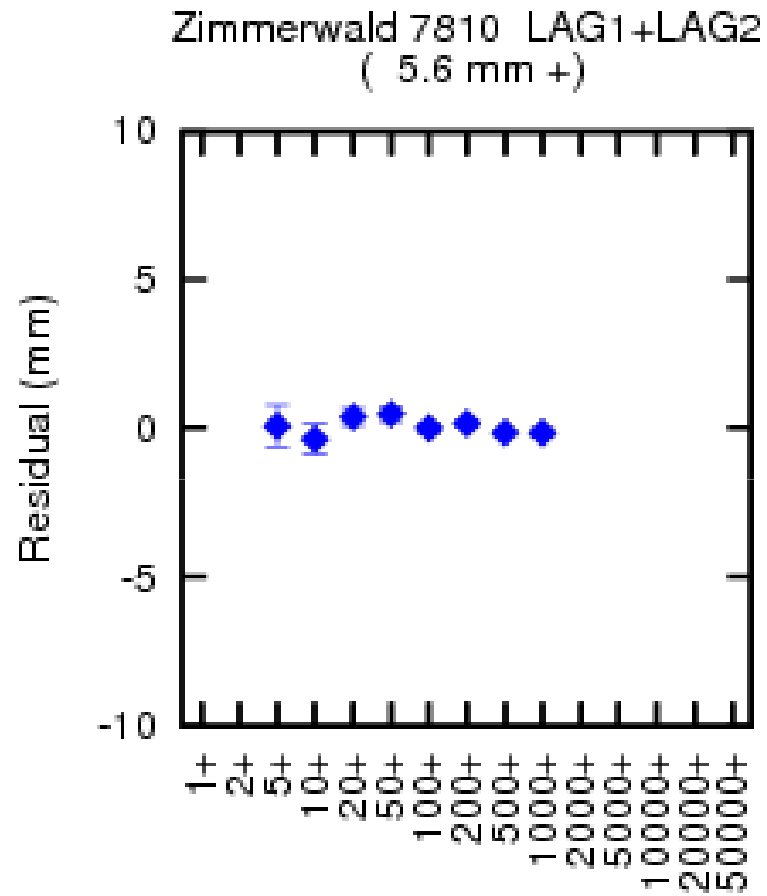
Yarragadee 7090 LAG1+LAG2
(0.9 mm +)



Yarragadee 7090 AJI
(14.2 mm +)

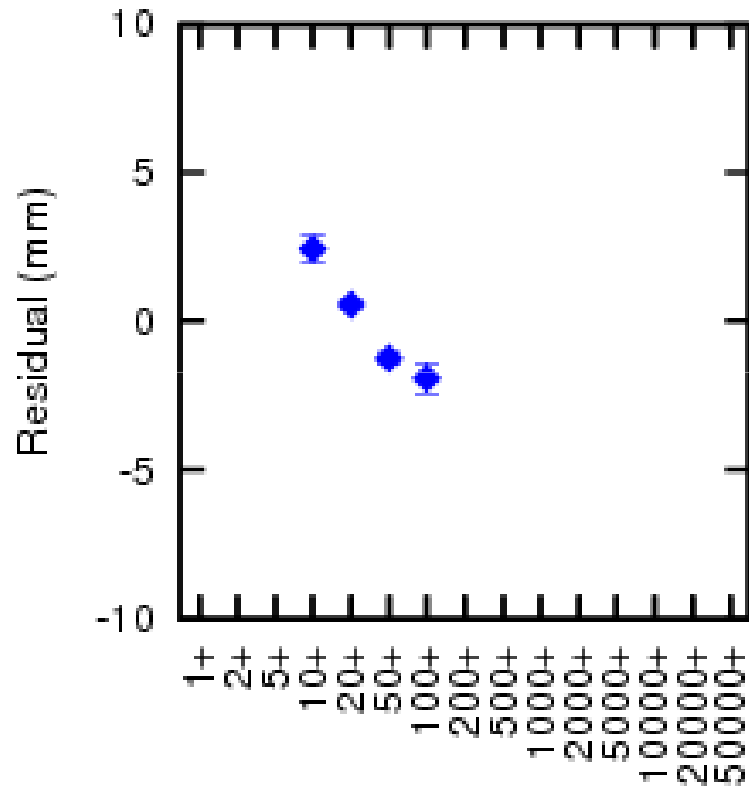


Bias vs Intensity 2010

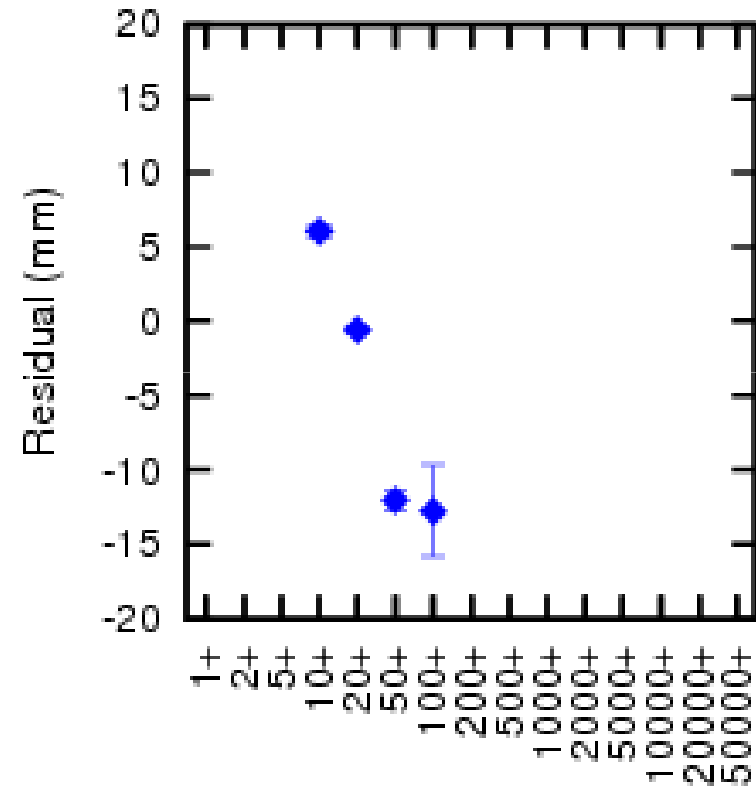


Bias vs Intensity 2010

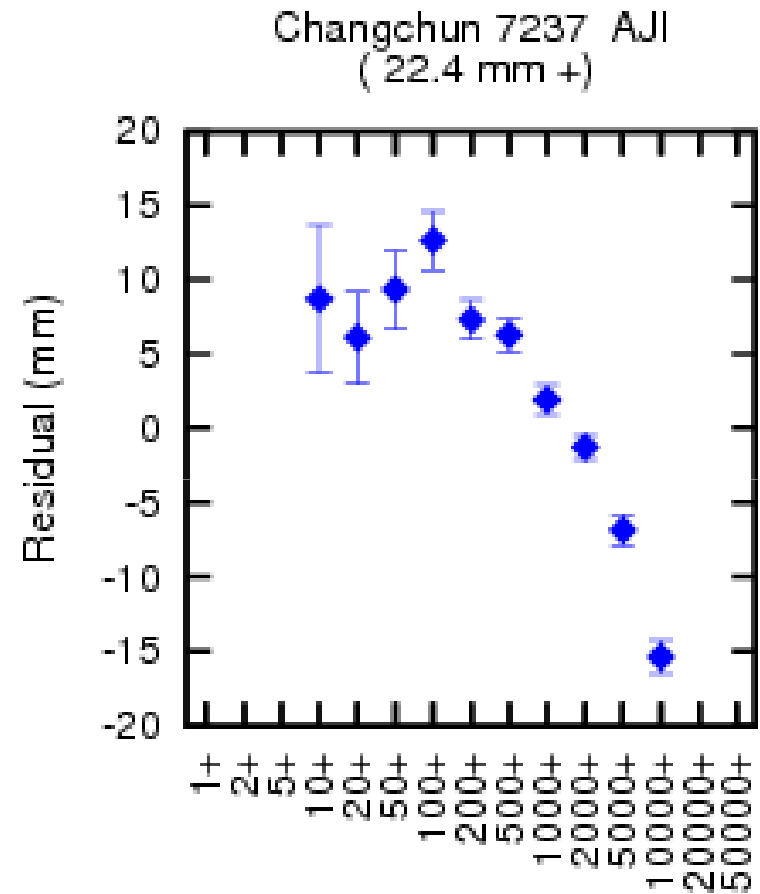
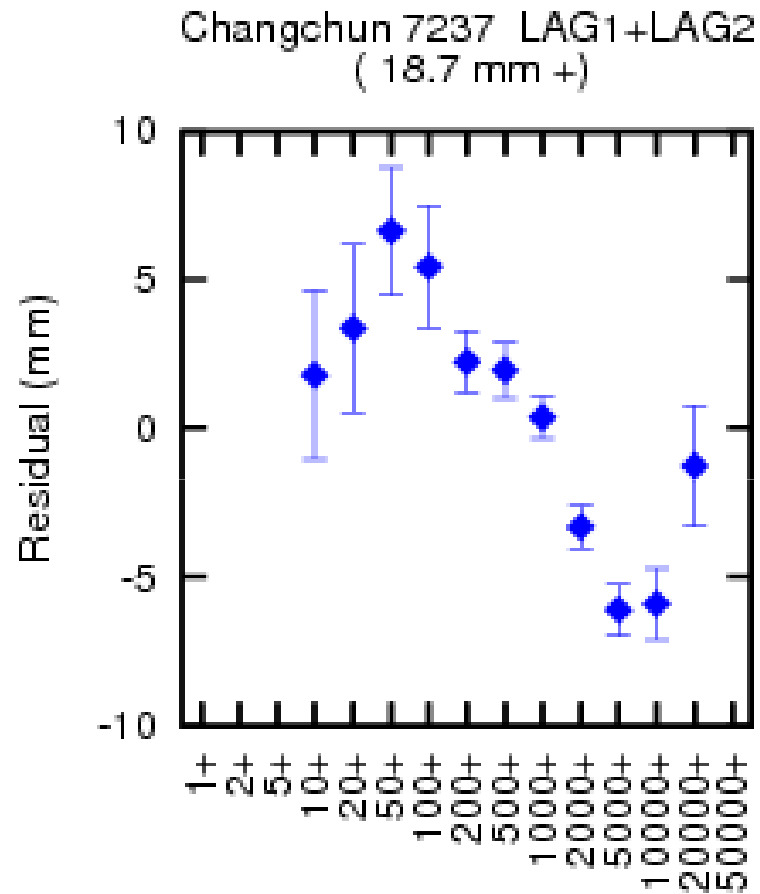
Mt Stromlo 7825 LAG1+LAG2
(2.5 mm +)



Mt Stromlo 7825 AJI
(8.2 mm +)



Bias vs Intensity 2010



**Contact me this week, or visit
<http://geo.science.hit-u.ac.jp/>
next week, for more stations
(up to most productive 20).**

Werner Gurtner says:

- [1] The laser tracking stations are continuously spending huge amount of budget (compared to the analysis institutes, and also to the GNSS stations).**

- [2] The tracking stations can not guarantee mm-level accuracy.**

[First half of this session]

How can the ACs handle the bias-hidden data?

How can the ACs make the best use of the current tracking data for their best products?

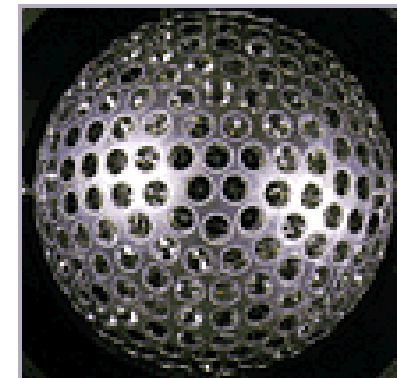
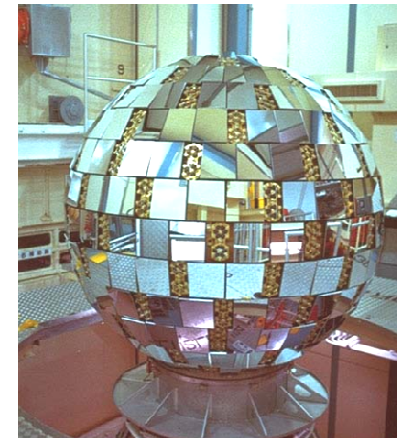
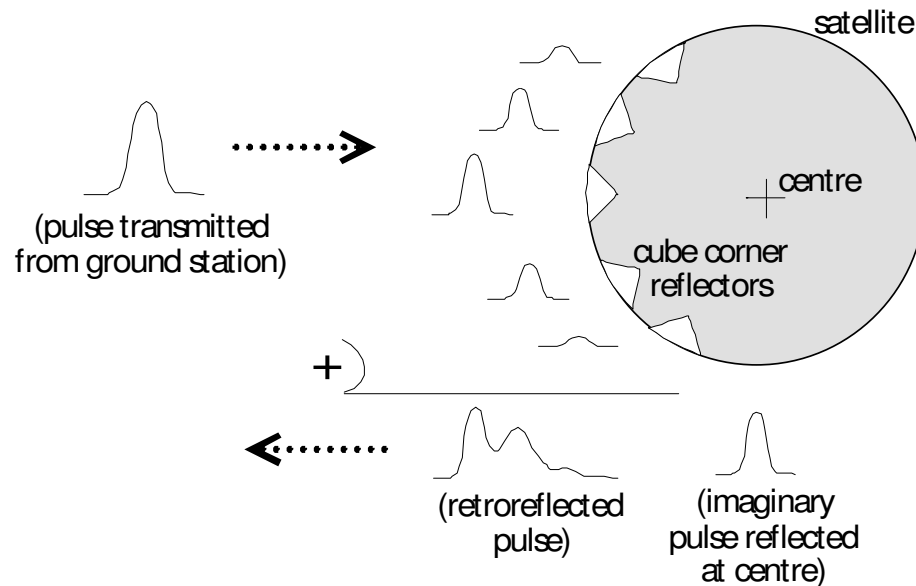
Spin: “Spin-off” of SLR

SLR can measure the satellite spins as well as the orbits.

Ajisai: Otsubo, IEEE TGARS, 2000.

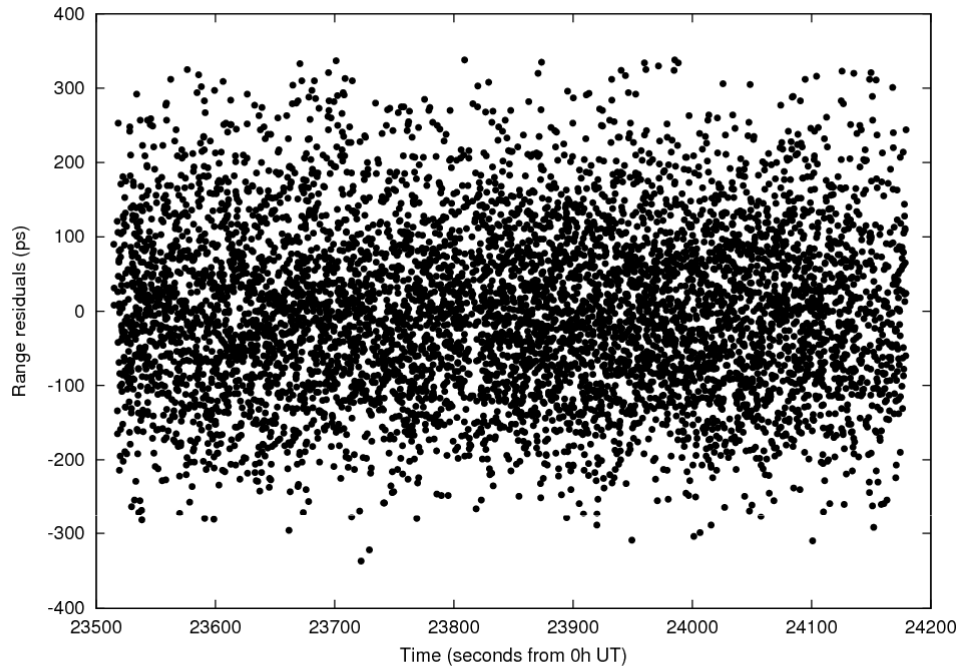
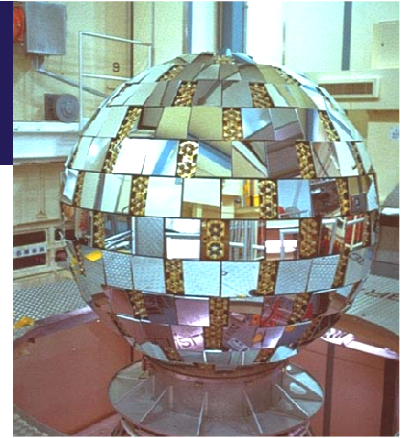
Lageos-2: Bianco, JGR, 2001.

Various: Kirchner/Kucharski, 200x-201x.



Less affected by SLR systematic error, utilising its high “precision”.
Precise spin parameters can improve the orbit determination.

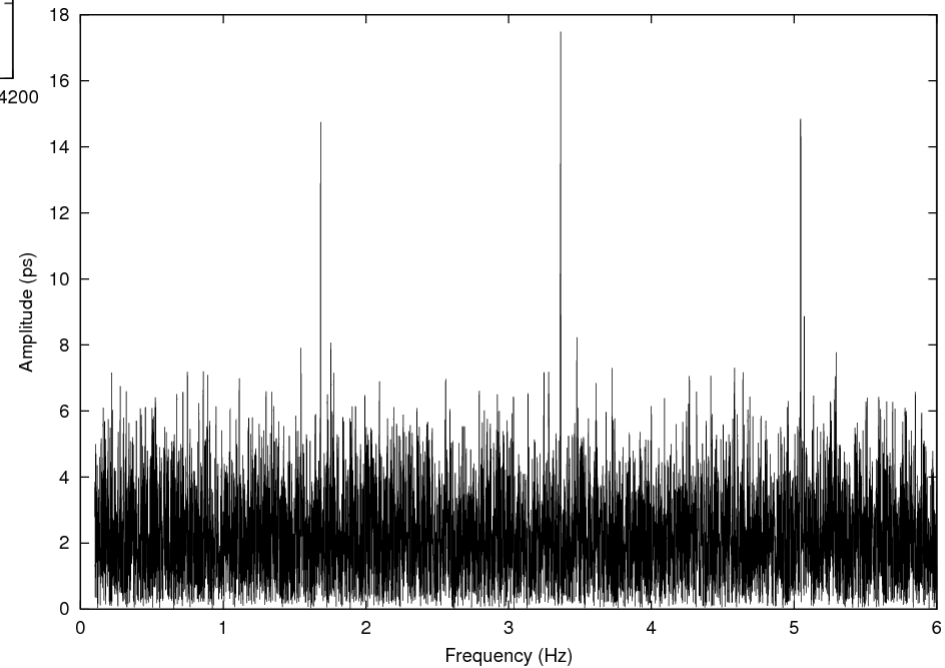
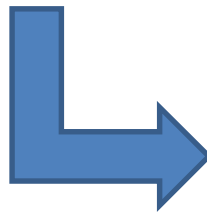
(Old) Example of Ajisai spin



→ Spin rate/period

3rd harmonics 6th harmonics 9th harmonics

Spectral Analysis

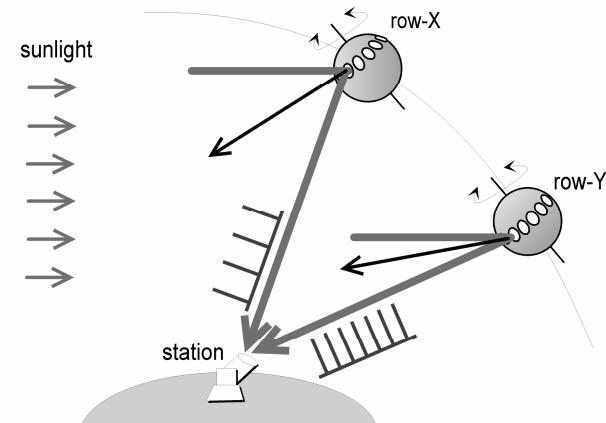


Otsubo, IEEE TGARS, 2000.

Methods to obtain the satellite spin/orientation

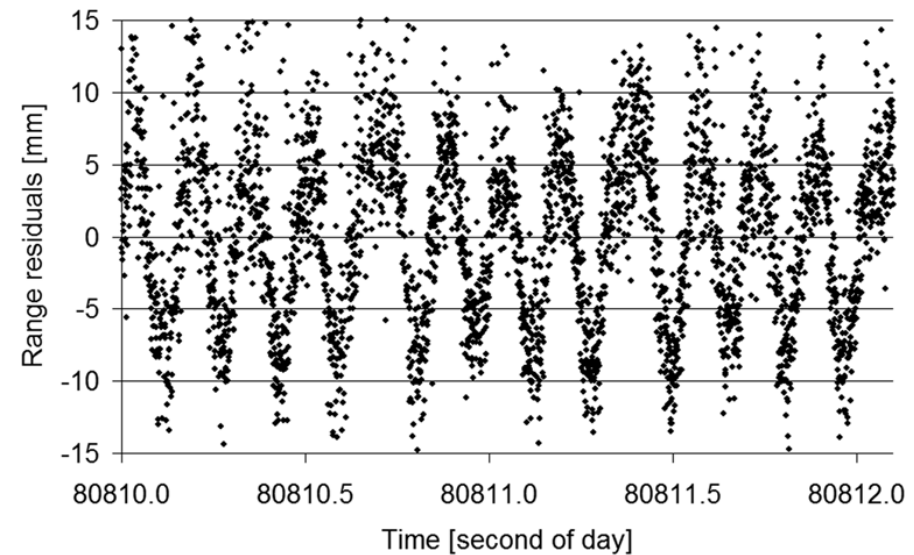
[Second half of this session]

[1] Photometry



Otsubo, IEEE TGARS, 2004.

[2] SLR residuals



Kucharski, JPGU, 2009.

[3] POD-based, More?