

# Optical Response Simulation for ASTRO-G Laser Reflector Array

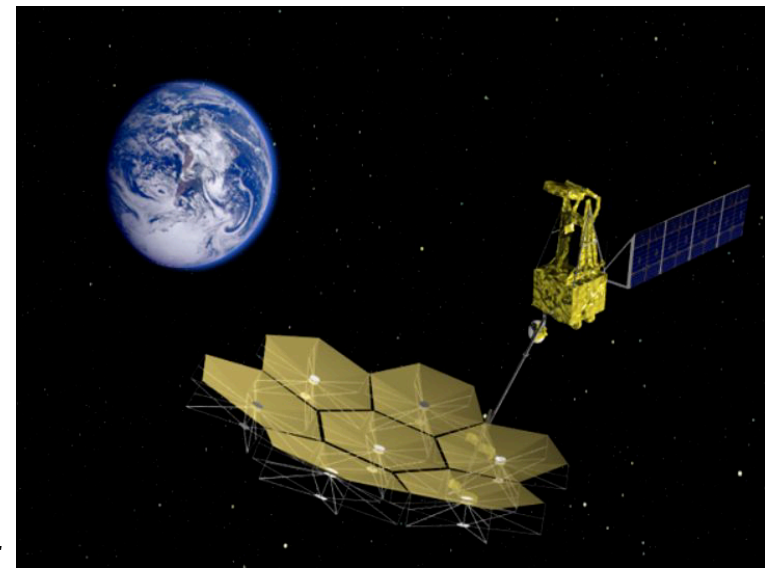
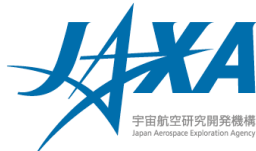
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# ASTRO-G Highly Elliptic Orbit

## ASTRO-G (VSOP-2)

Launch: FY2012 (5-year mission)

9.6-metre antenna

Highly elliptic orbit: 1000 x 25000 km altitude

Observation bands: 8.4, 22 and 43 GHz

High frequency, High resolution and High sensitivity



## 4 x GPS (+Galileo?) Receiver

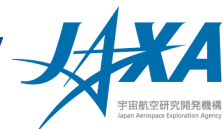
Effective <3000-5000 km only (1 hr per 7.5 hrs).

Sidelobe? One-frequency use?

## 1 x SLR Retroreflector Array ← This talk

Should be effective 1000 km to 25000 km

More than a cal/val instrument



# 4-D Simulation of CCR Response

## 4-Dimensional Function:

Angle of incidence and azimuth (2-D)

Velocity aberration (2-D)

## Software development for Single CCR Response (ongoing)

Language: C#

Input:

CCR Shape, Optical Index, Coat, Size, Recession, Dihedral angle

Laser wavelength, Polarisation

Output:

Far-field amplitude

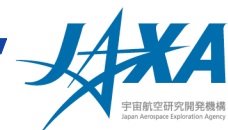
Grid size: 2-deg for angle of incidence, 2- $\mu$ rad for velocity aberration

> 2 GB in ASCII Text, > 100 MB in Binary (NetCDF) file

Computation time: 6 to 14 hours per reflector ... needs optimisation



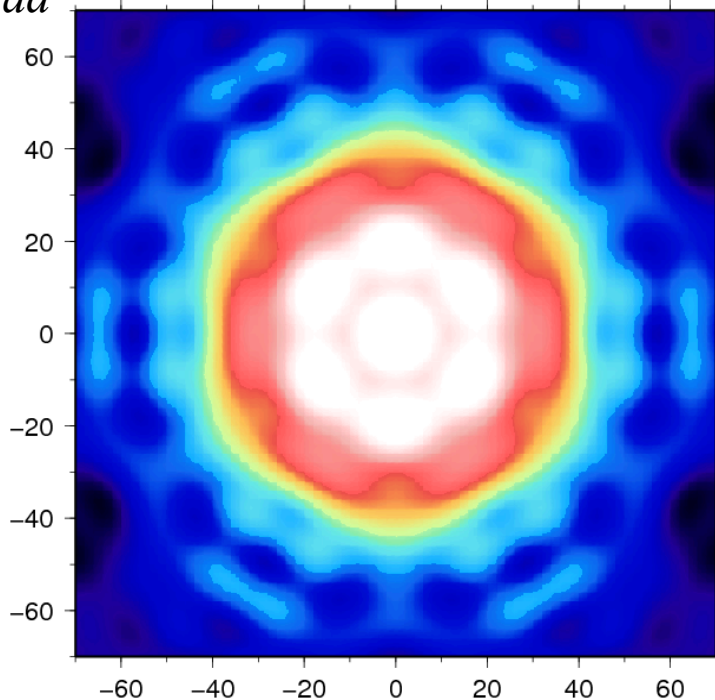
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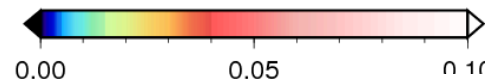
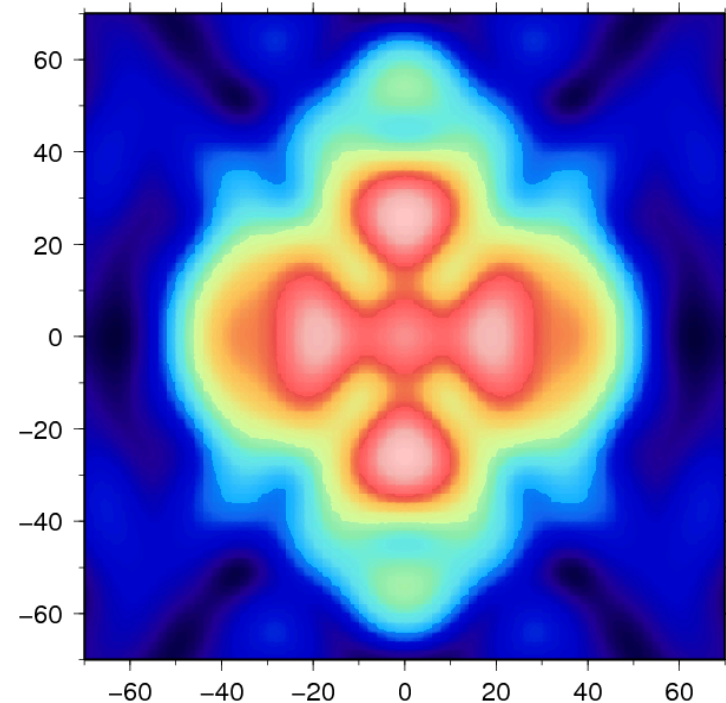
# Far Field Diffraction Pattern (examples)

38 mm uncoated CCRs, Dihedral angle = 0.75"  
Circular polarisation

$\mu$  rad Angle of incidence:  $i = 0$  deg,  $az = 0$  deg

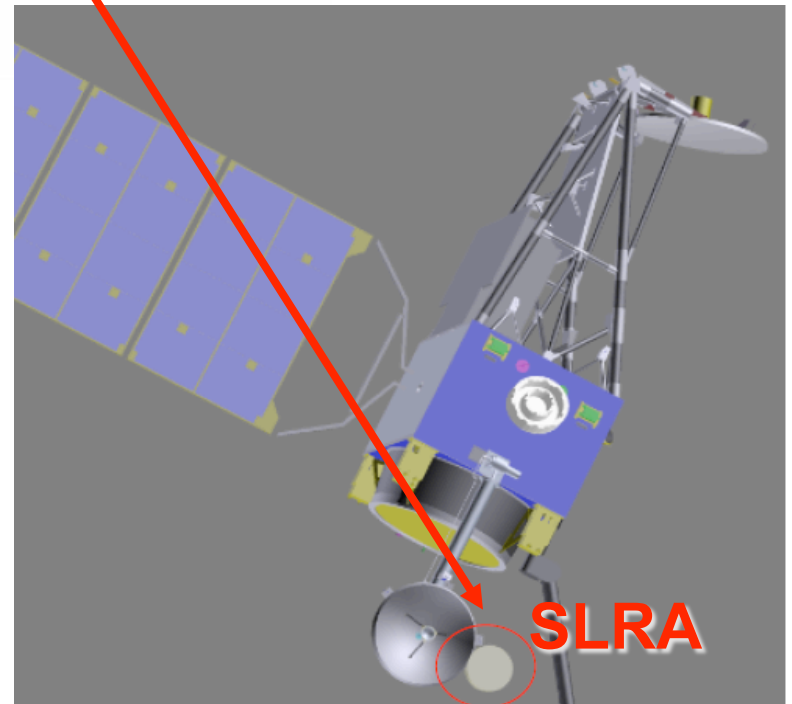
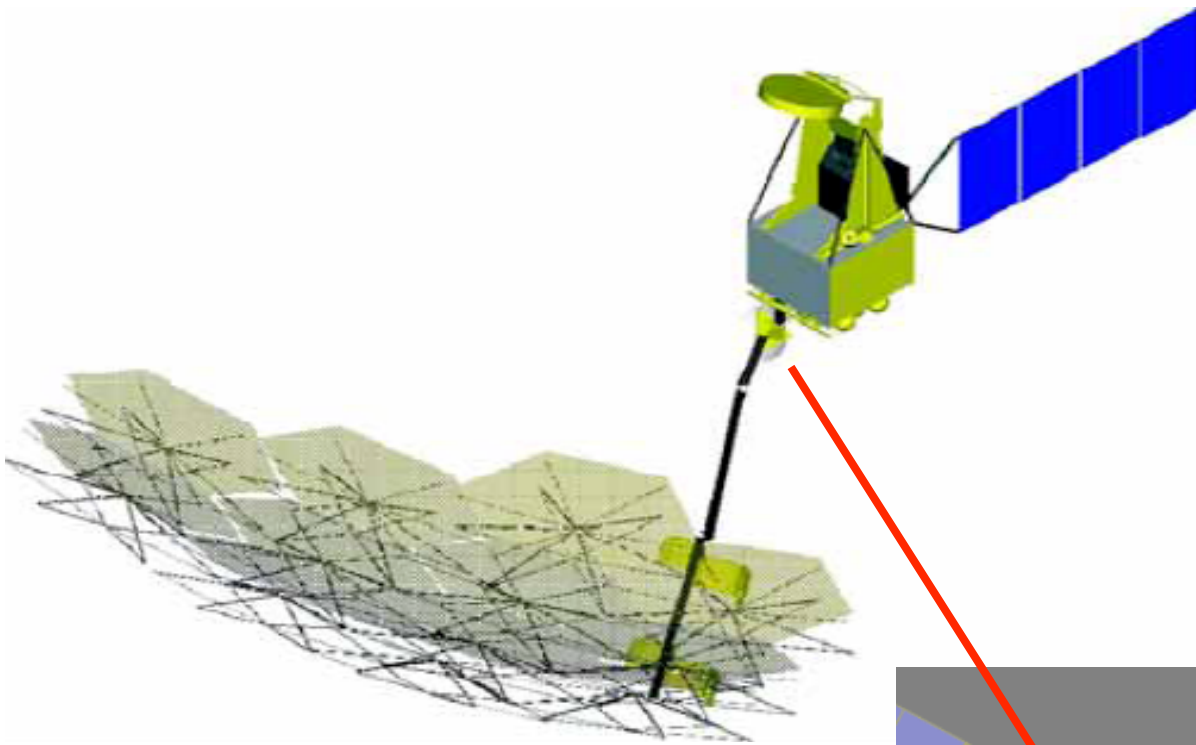


Angle of incidence:  $i = 16$  deg,  $az = 0$  deg



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# Problem 1: Angle of Incidence

## Orientation of ASTRO-G

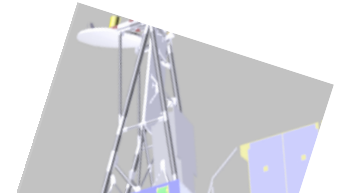
9.6-m antenna points stars  
by changing the satellite attitude

Ka antenna points one of Ka stars  
only when it is visible ( $> 5$  deg)  
above 5500 km

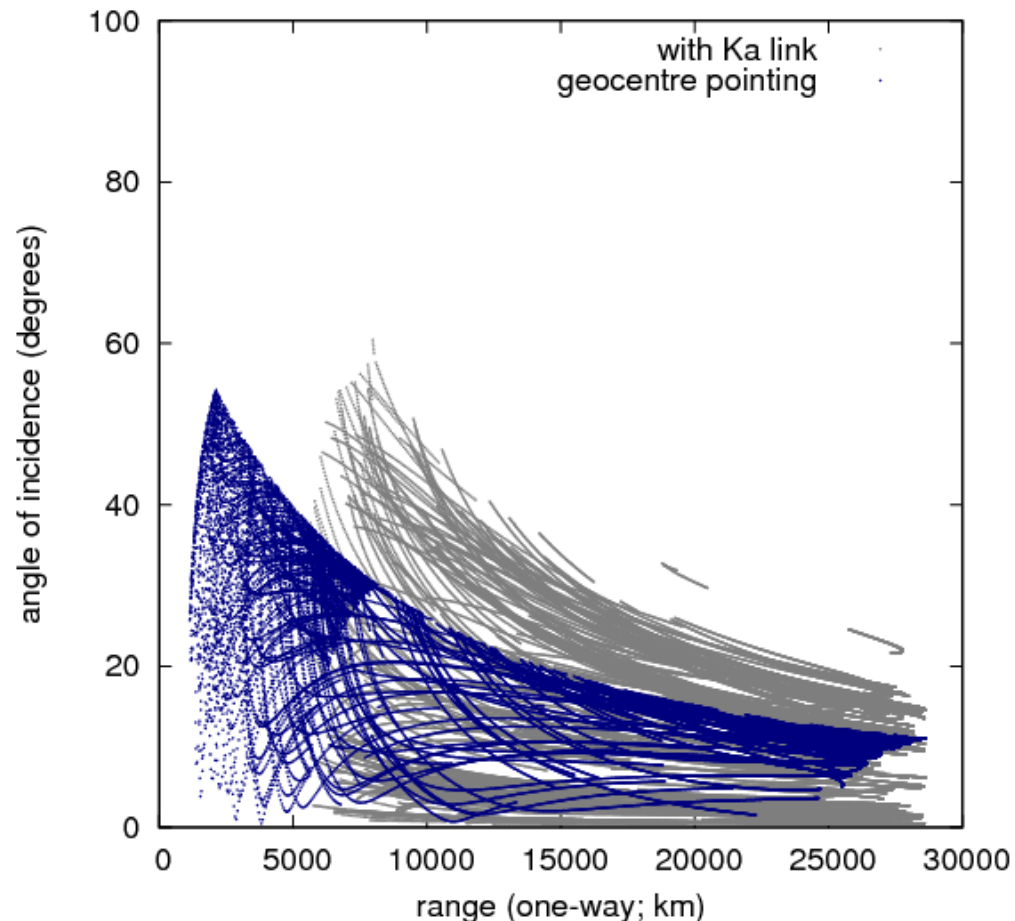
Otherwise,

Ka antenna points the geocenter

SLR array sync. with Ka antenna

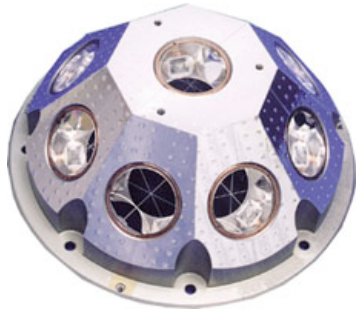


ASTRO-G (3 Ka stations: active  $> 5500$  km alt)





# Reflector Array for LEO and GNSS



**ALOS**  
**720 km circular orbit**

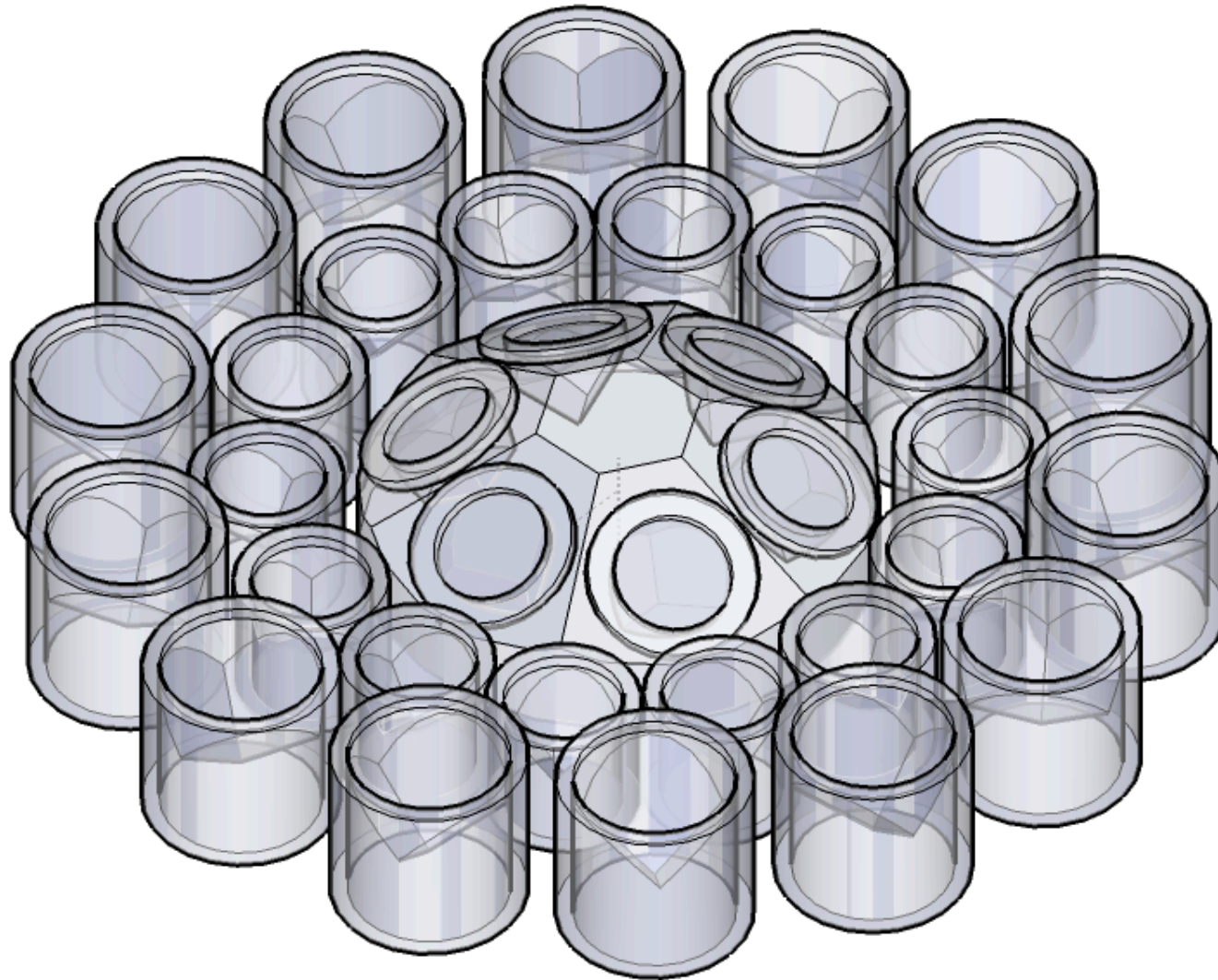


**GPS**  
**20000 km circular orbit**

Retro pictures: © ILRS Web



# Reflector Array: Basic Design



## Centre:

6 x

28 mm coated CCRs

Slanted by 30 deg

## Inner Ring:

14 x

28 mm uncoated CCRs

## Outer Ring:

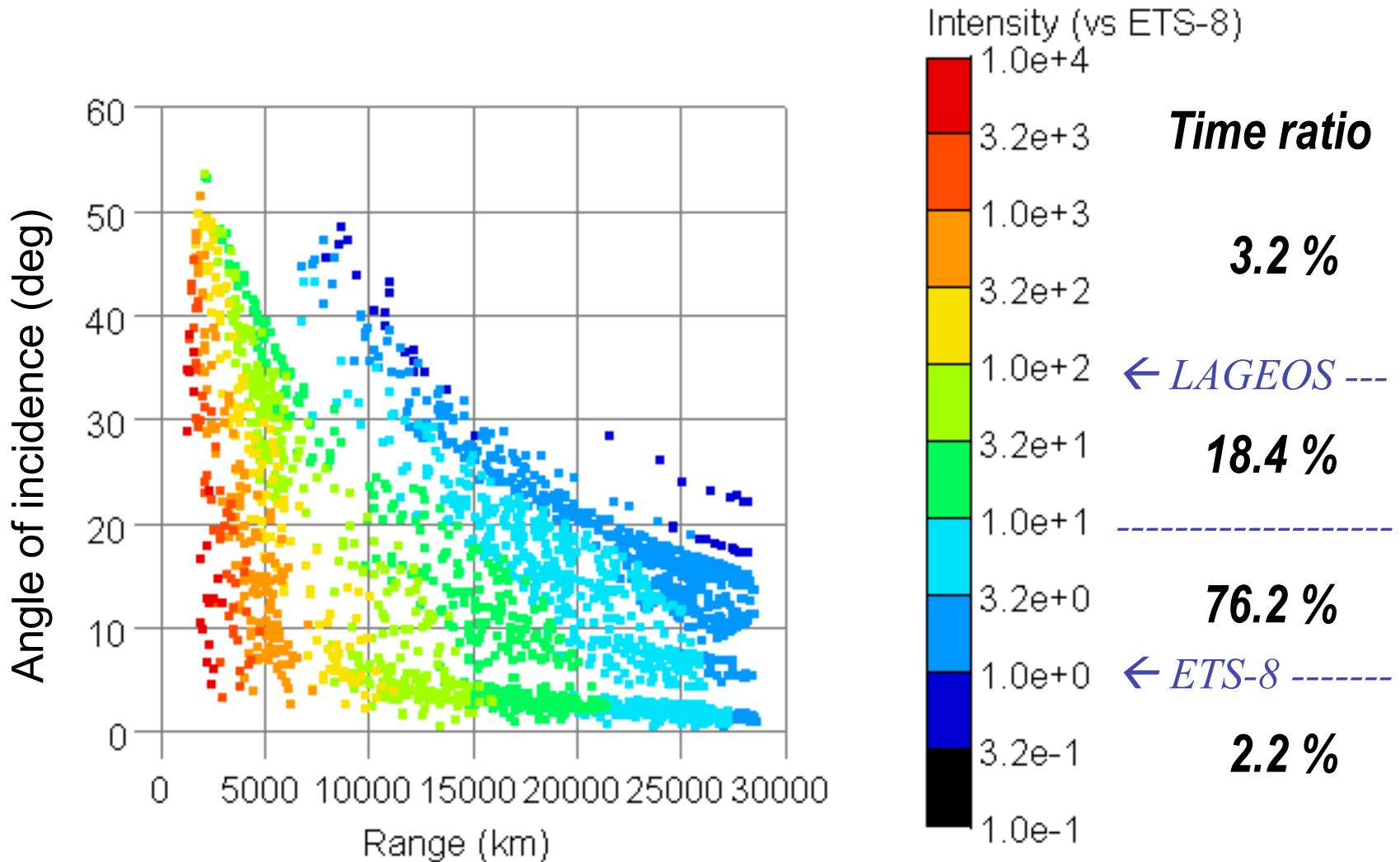
14 x

38 mm uncoated CCRs





# Intensity: ASTRO-G vs ETS-8

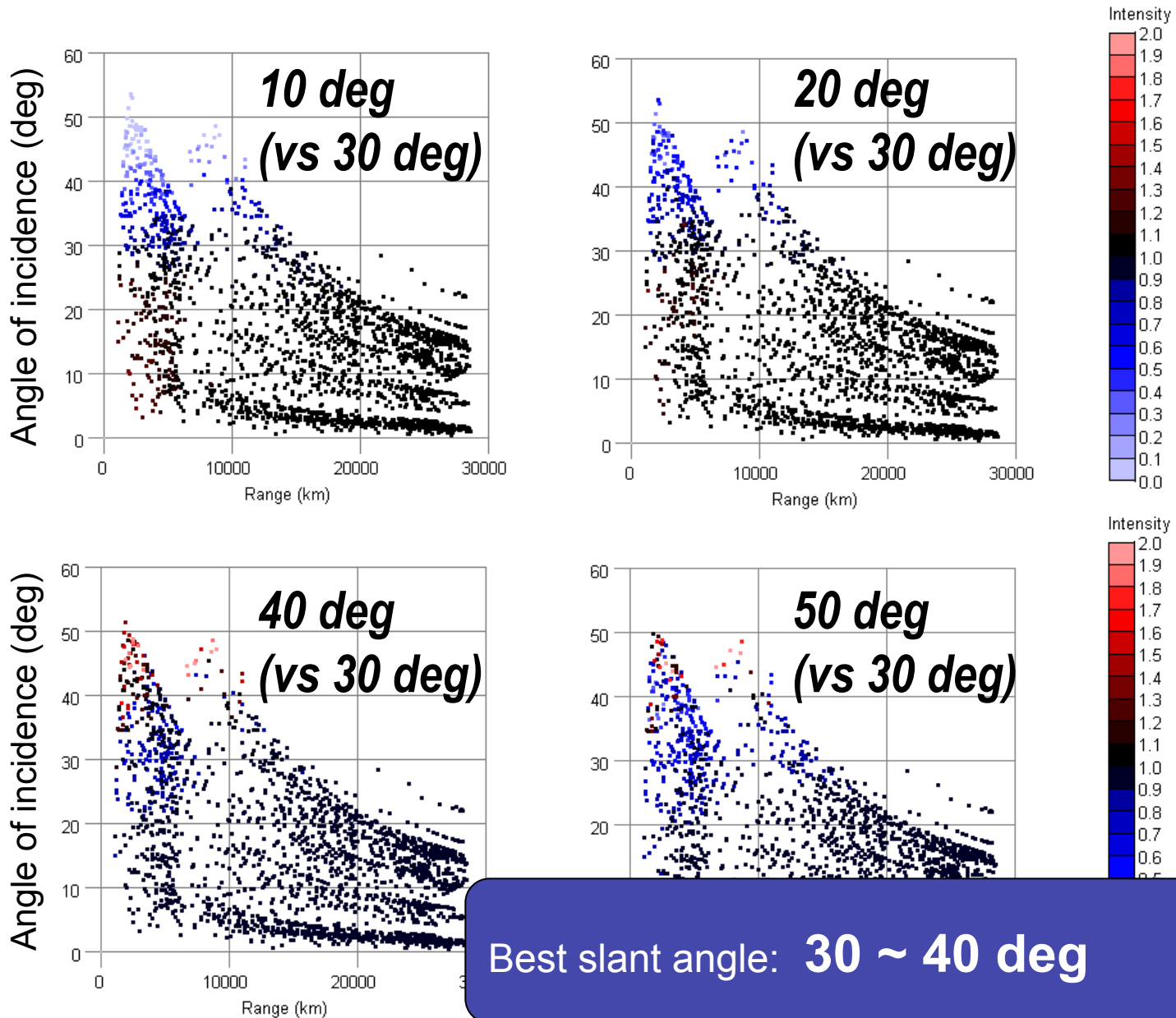


ETS-8 intensity: assuming a 30.5 N station (Tanegashima)



# Slant angle of centre reflectors

Nominal  
30 deg



Best slant angle: 30 ~ 40 deg

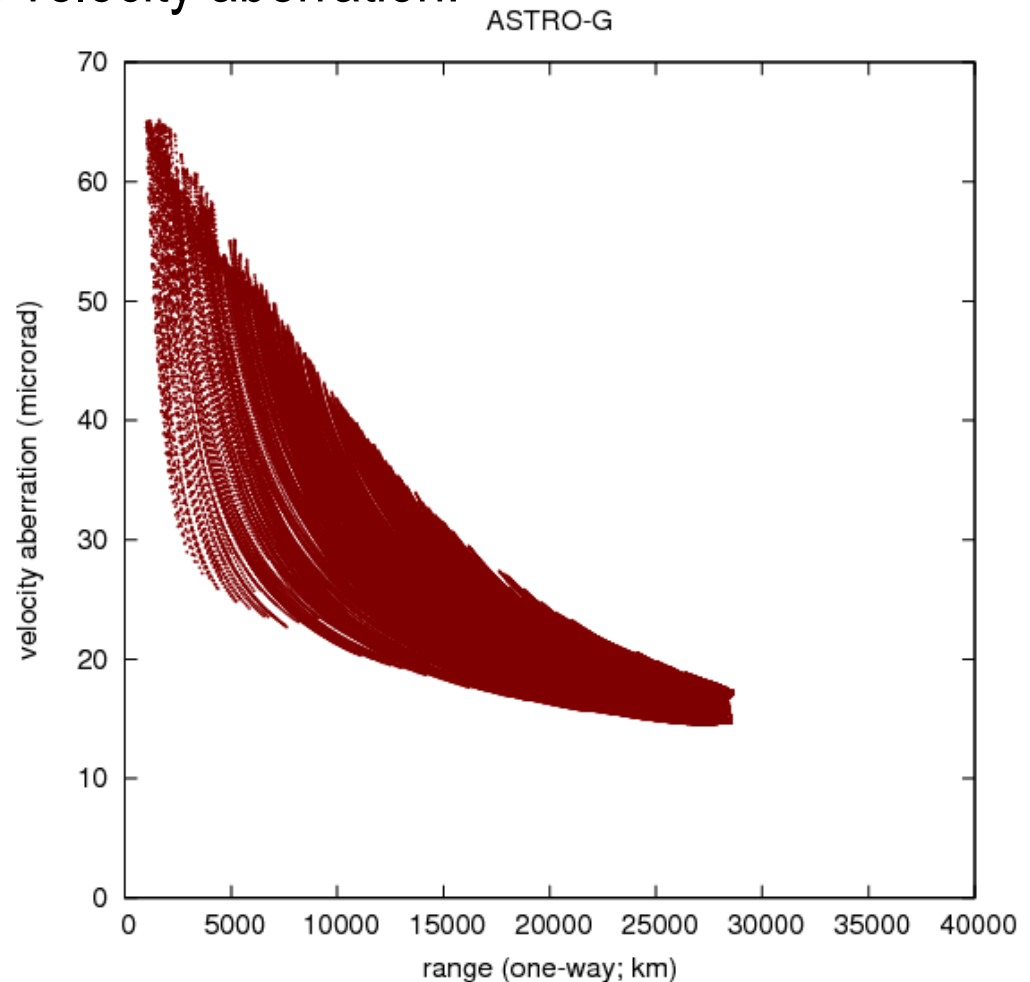
# Problem 2: Velocity Aberration

## Large Velocity Variation of ASTRO-G

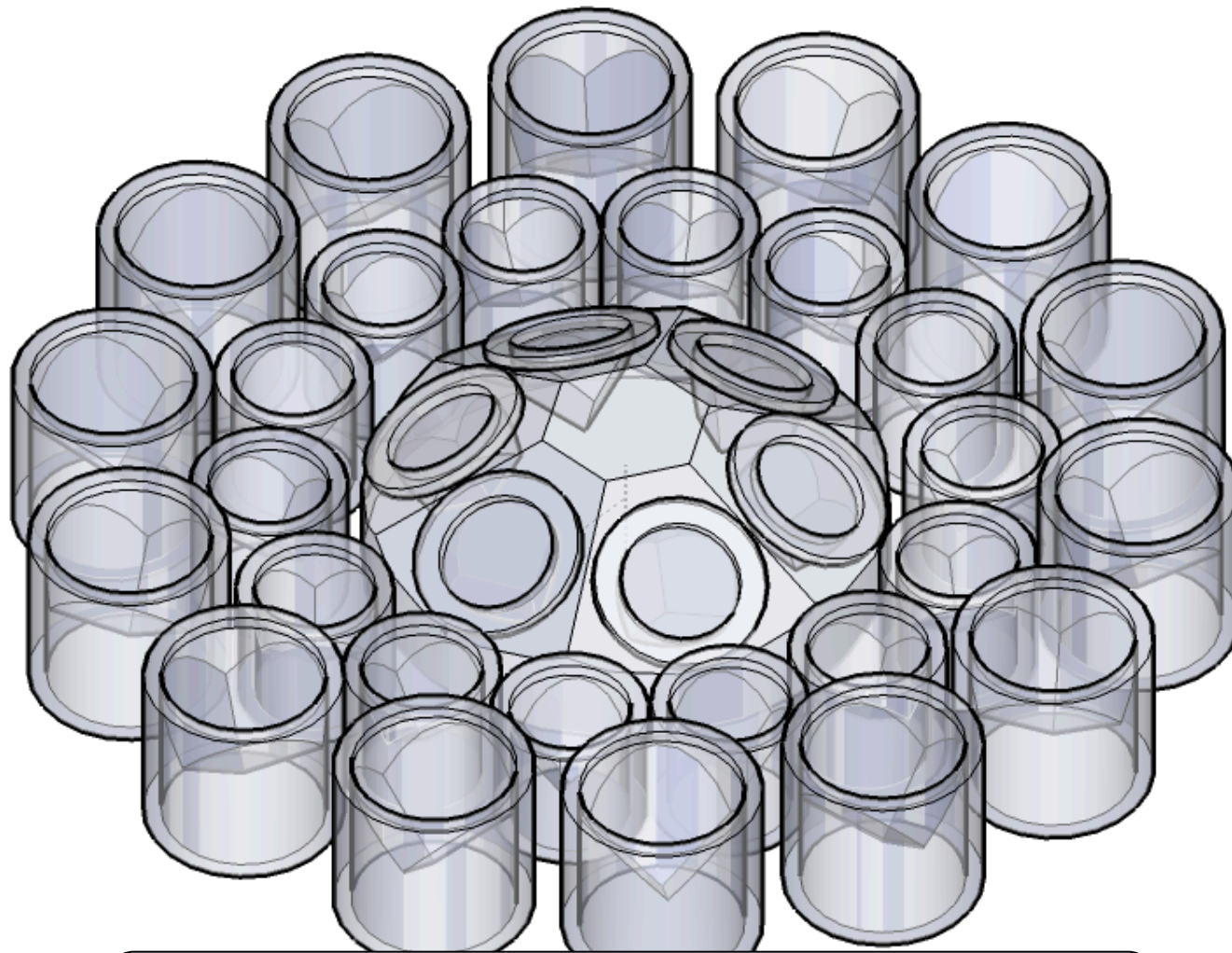
2.2 km/s (apogee) to 9.4 km/s (perigee)

Unprecedented wide range of velocity aberration!

15 to 65 microrad



# Reflector Array: Dihedral Angle



**Centre:** Dihedral angle  
~ **2.0''**  
6 x  
28 mm coated CCRs  
Inclined by 30 deg

**Inner Ring:** Dihedral angle  
< **0.75''**  
14 x  
28 mm uncoated CCRs

**Outer Ring:** Dihedral angle  
< **0.75''**  
14 x  
38 mm uncoated CCRs

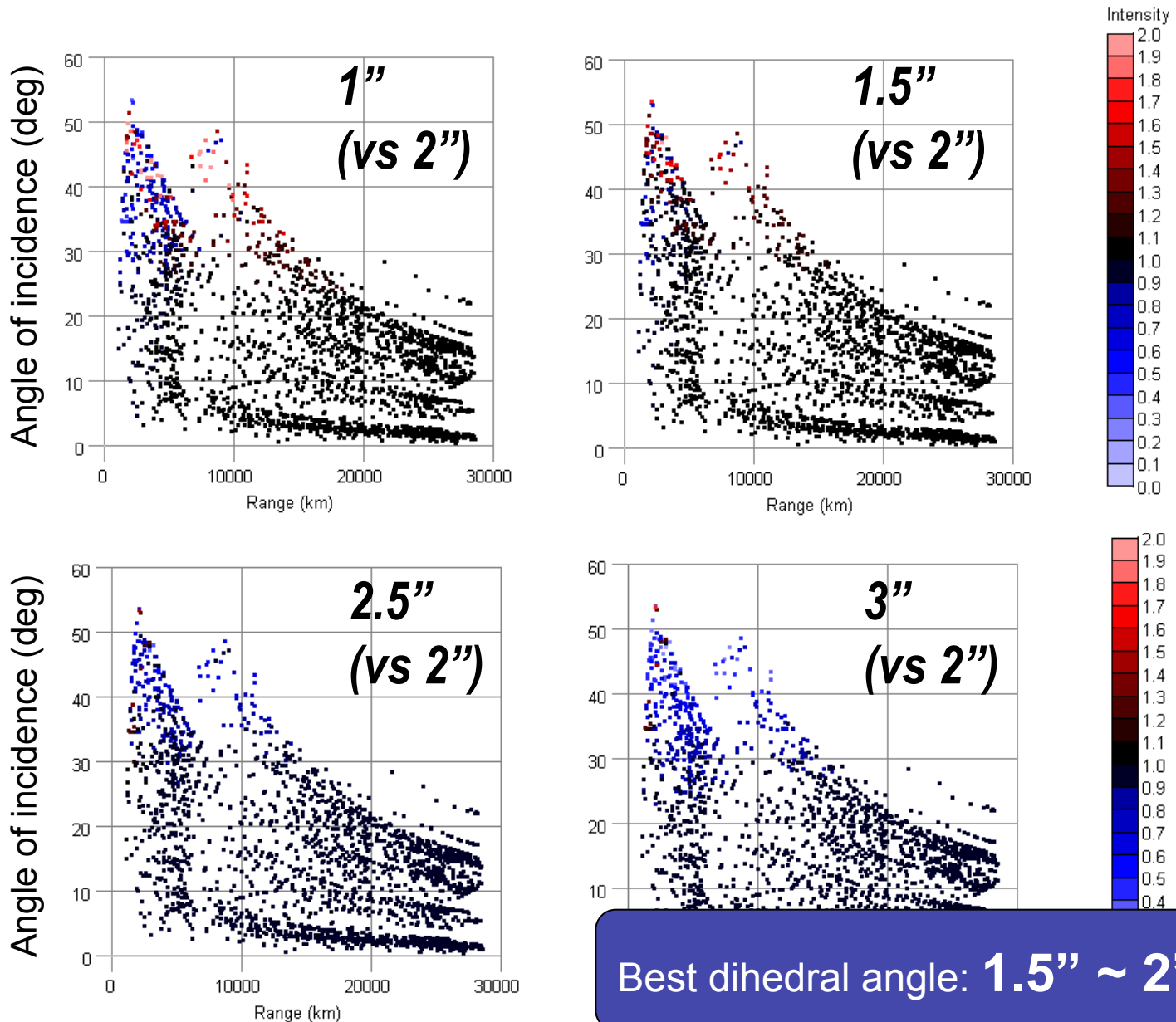
cf. Dihedral angle

~ **1.25'' (LAGEOS)**  
~ **0.50'' (ETS-8)**



# Dihedral angle of centre reflectors

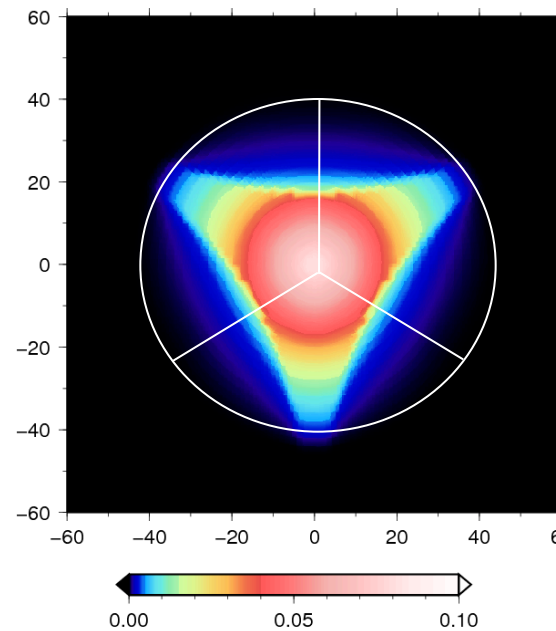
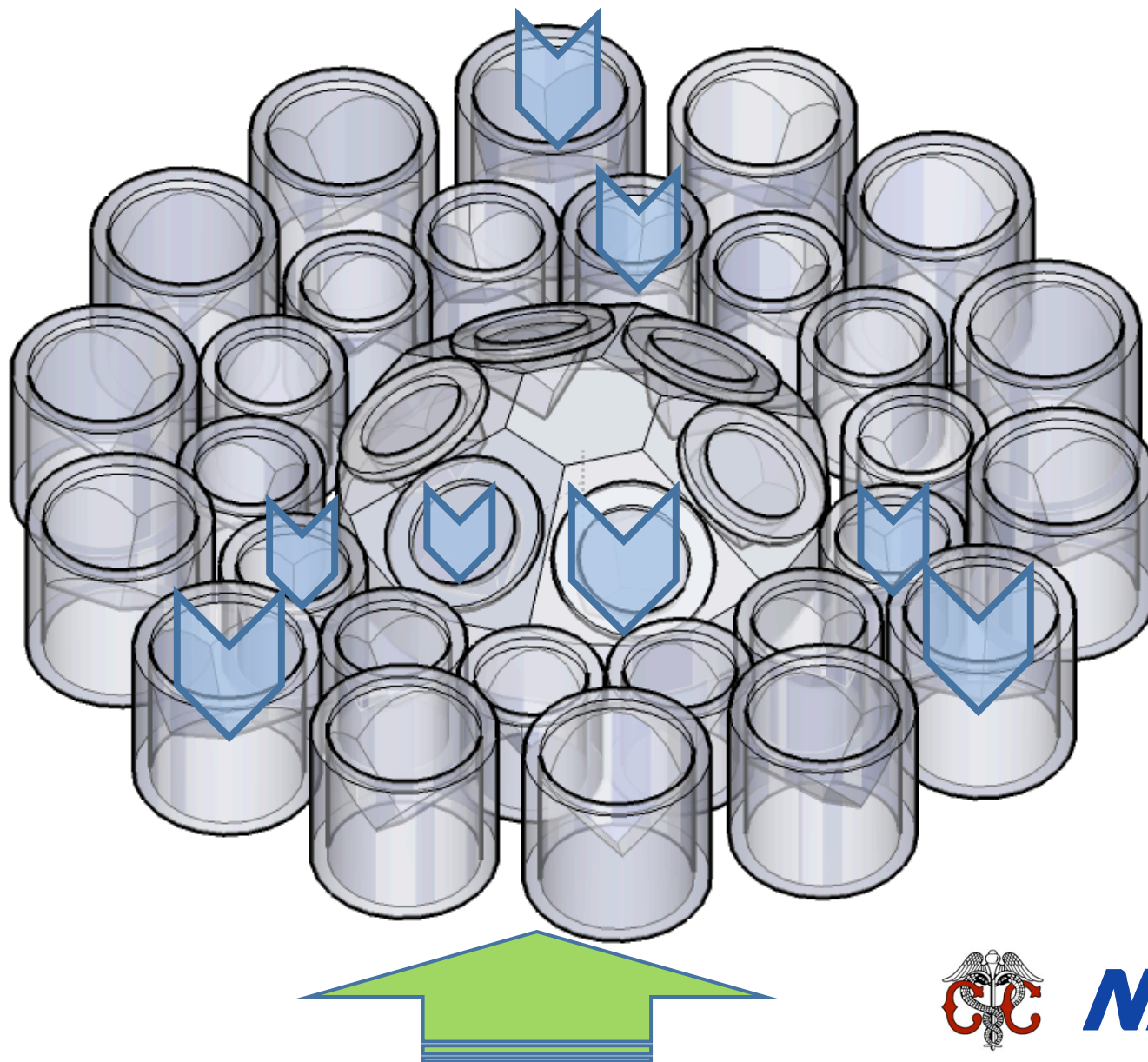
Nominal  
2.0"



Best dihedral angle: 1.5" ~ 2"



# Reflector Array: Orientation & Double Pulse



Single Reflector  
(uncoated)

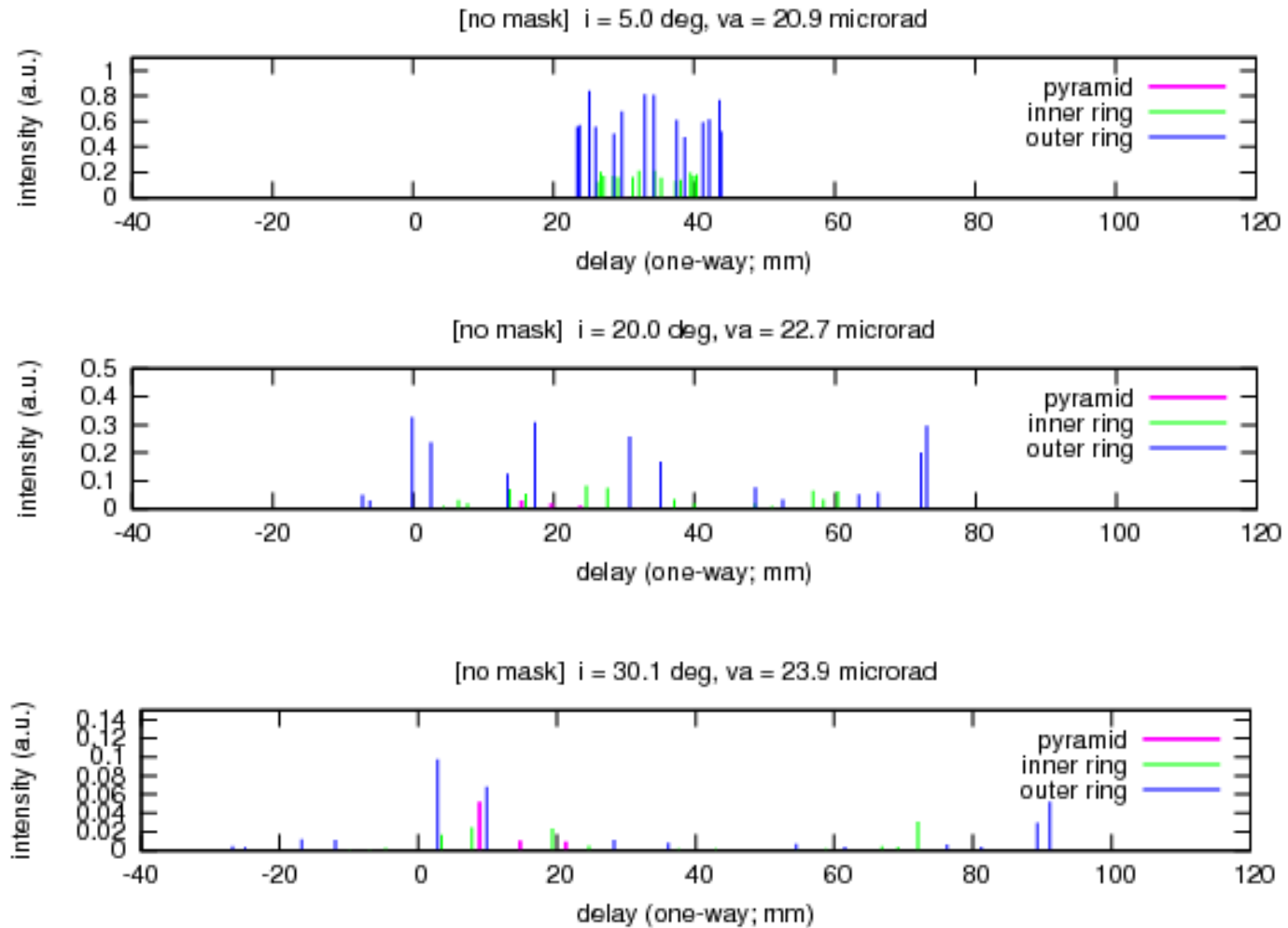
Strong 120-deg  
az-dependence



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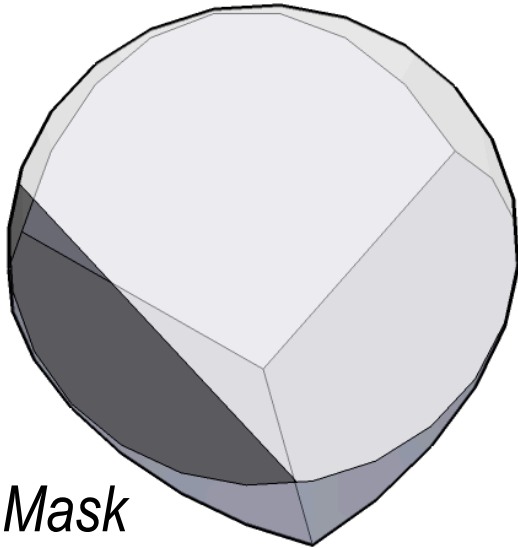


# Problem 3: Double pulse

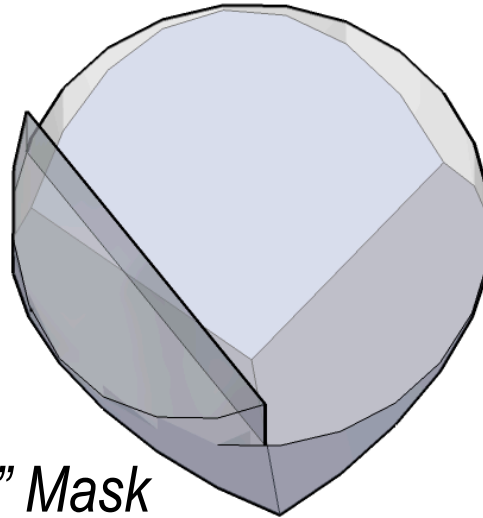


# Mask Ideas

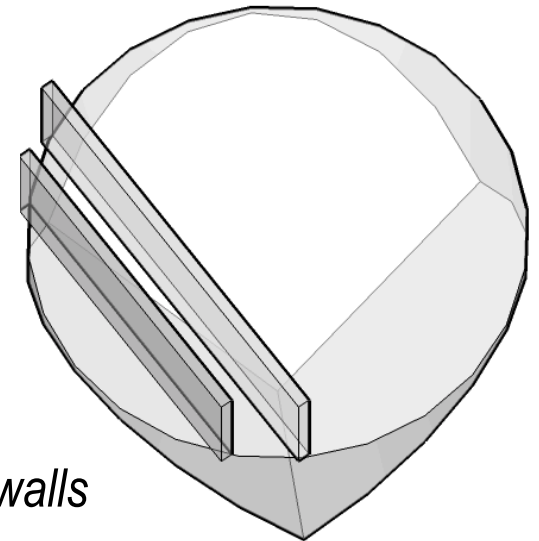
Precision  
vs  
Intensity



*“Flat” Mask*



*“Delta” Mask*



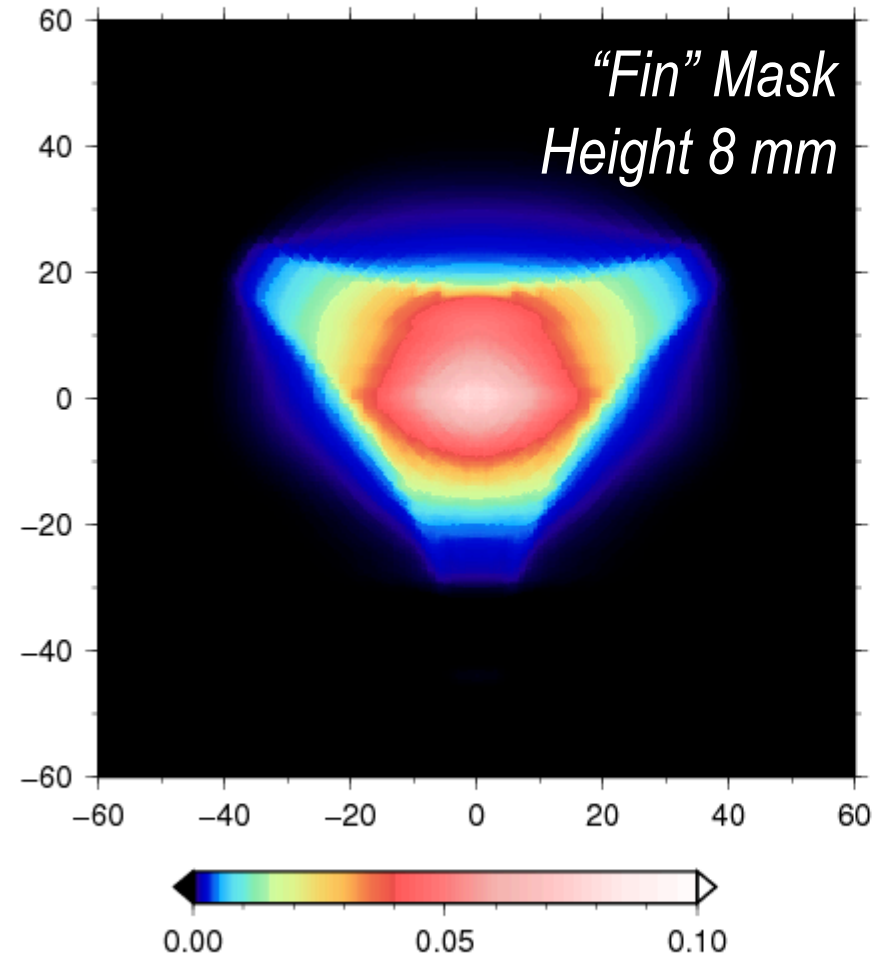
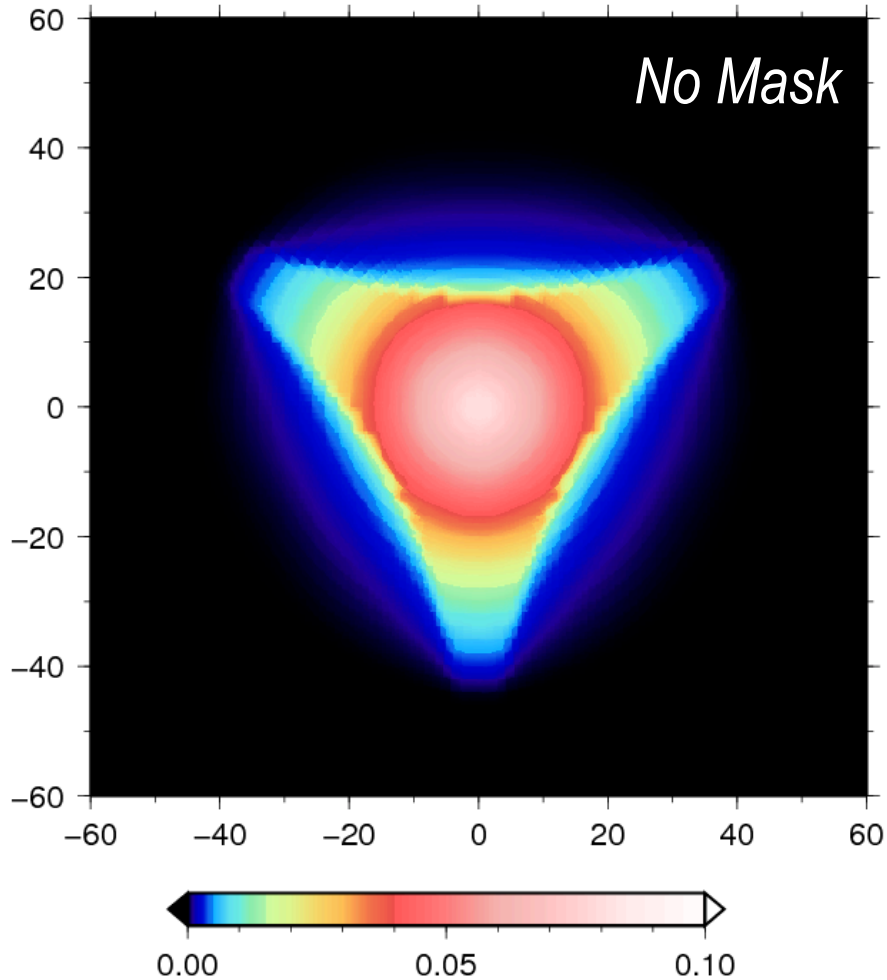
*“Fin” Mask*

*Two x-mm height walls  
at  $0.25 r$  and  $0.5 r$*



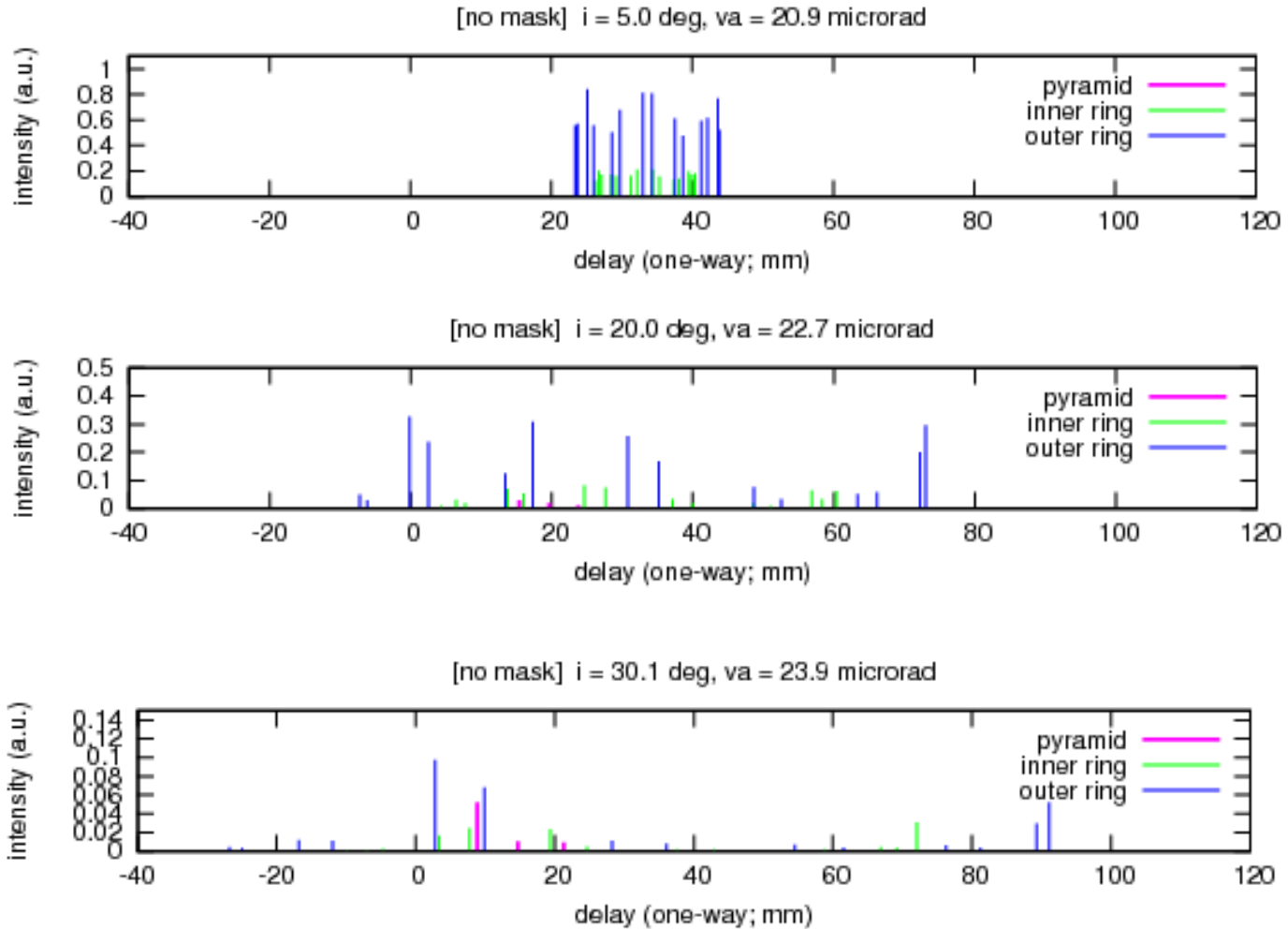
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# Effect of masks



# Problem 3: Double pulse

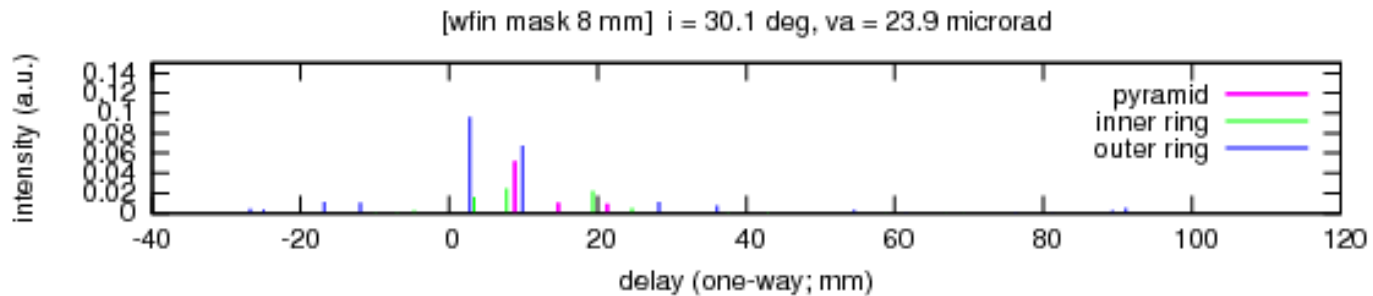
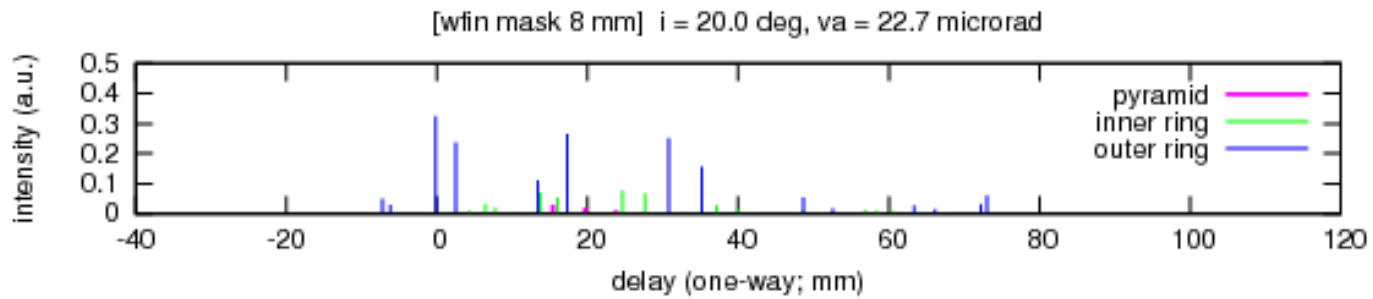
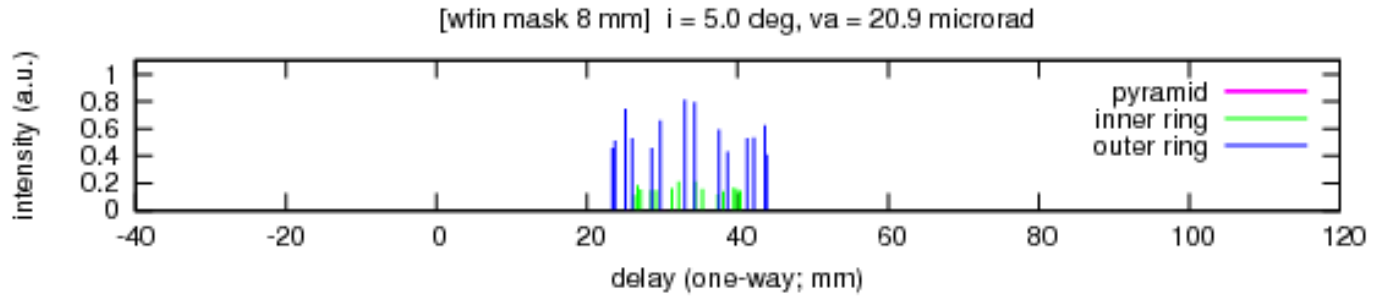
No  
Mask





# Double pulse eliminated!

Fin  
8 mm  
Height



# Summary

## Array design

Mixture of LEO-type and GNSS-type

Return Intensity: Mostly stronger than ETS-8, Huge variation in time

## Plenty of new problems/concepts due to its orbit

Angle of incidence

Velocity aberration

Double pulse elimination

## Basic design done, but still needs fine tuning

**We sincerely hope your supports to this challenging target.**

*Laser ranging helps blackhole studies for the first time ever!*

*(But its launch is FY2012 = still 4 years ahead)*



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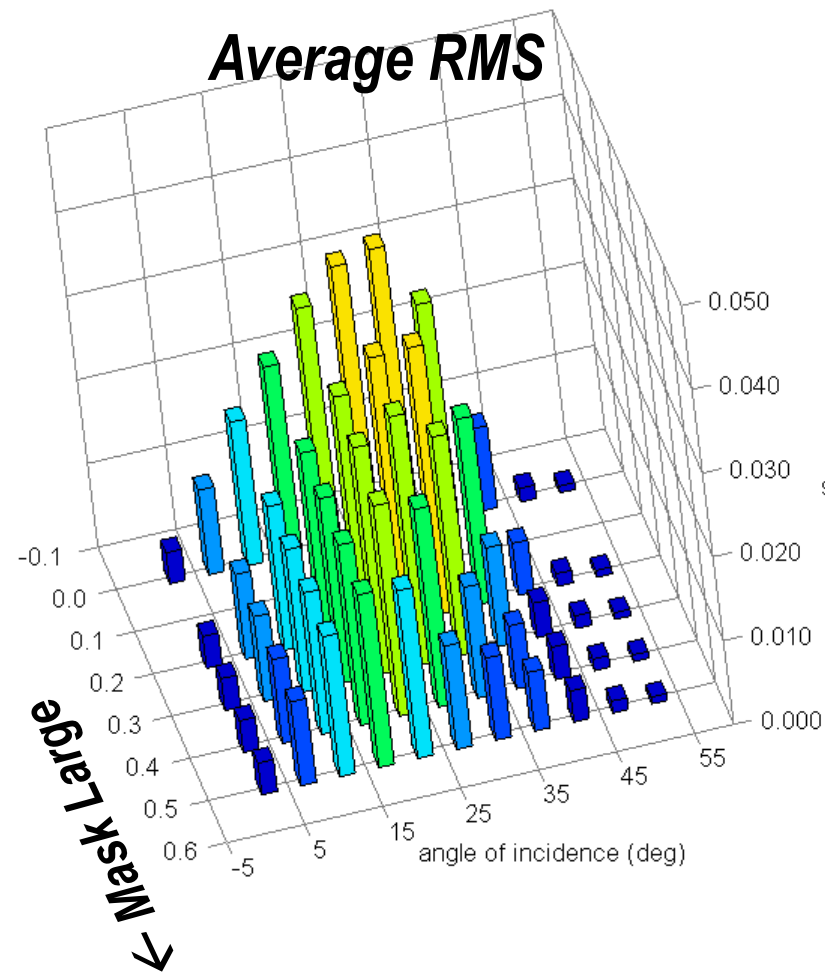




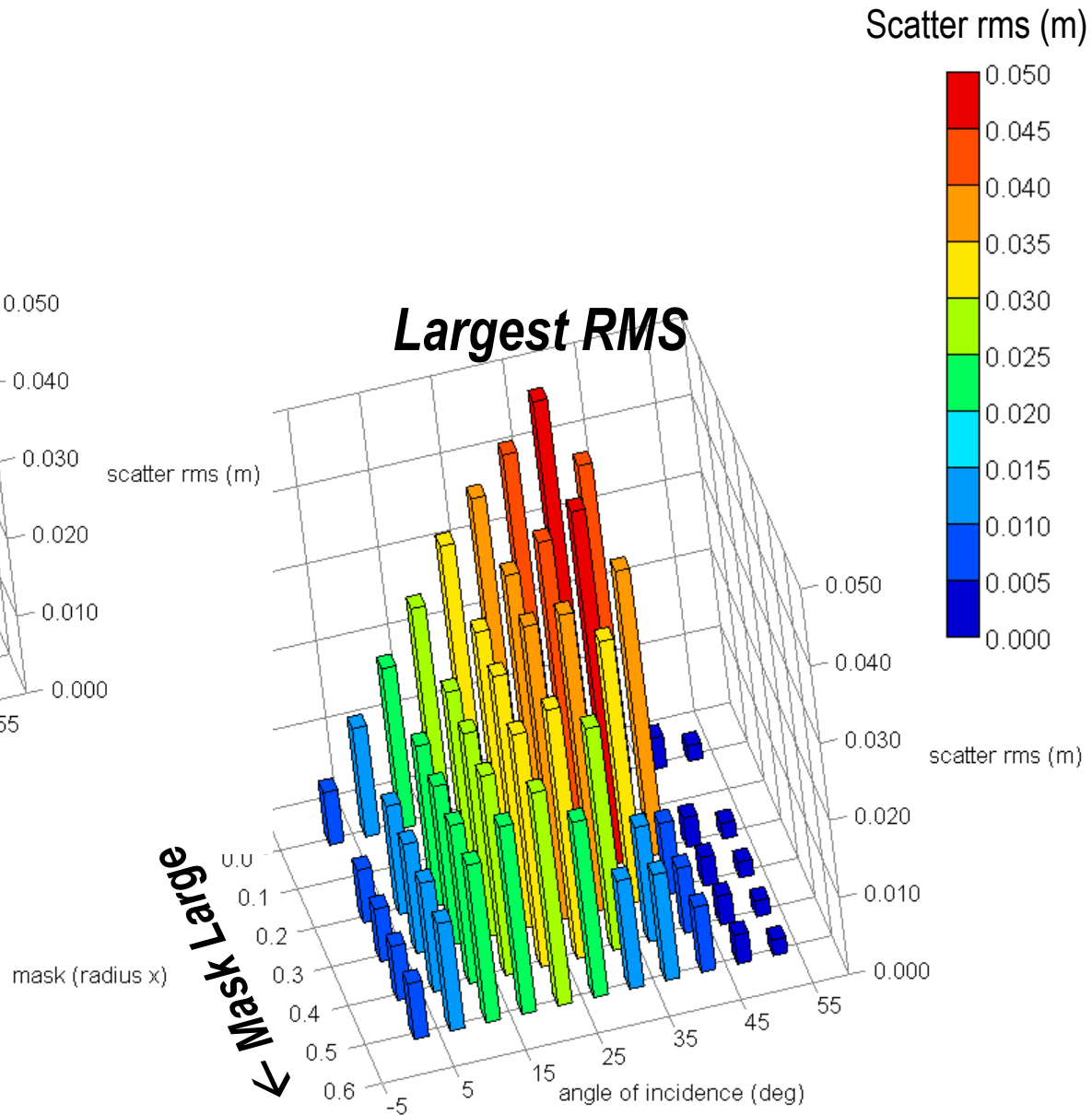


# Flat mask cases: Scatter RMS

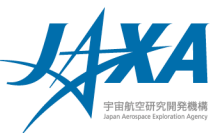
## Average RMS



## Largest RMS



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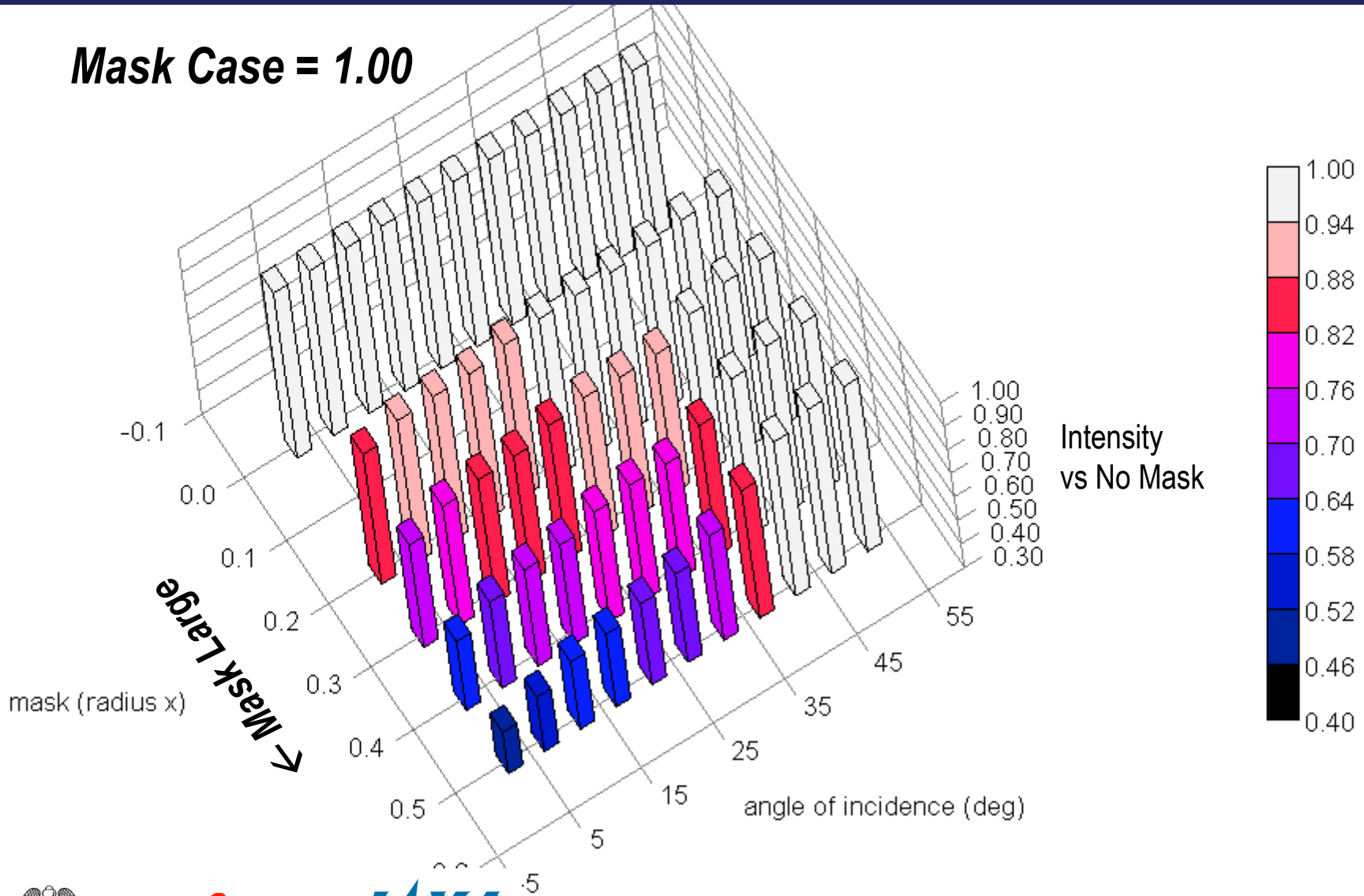


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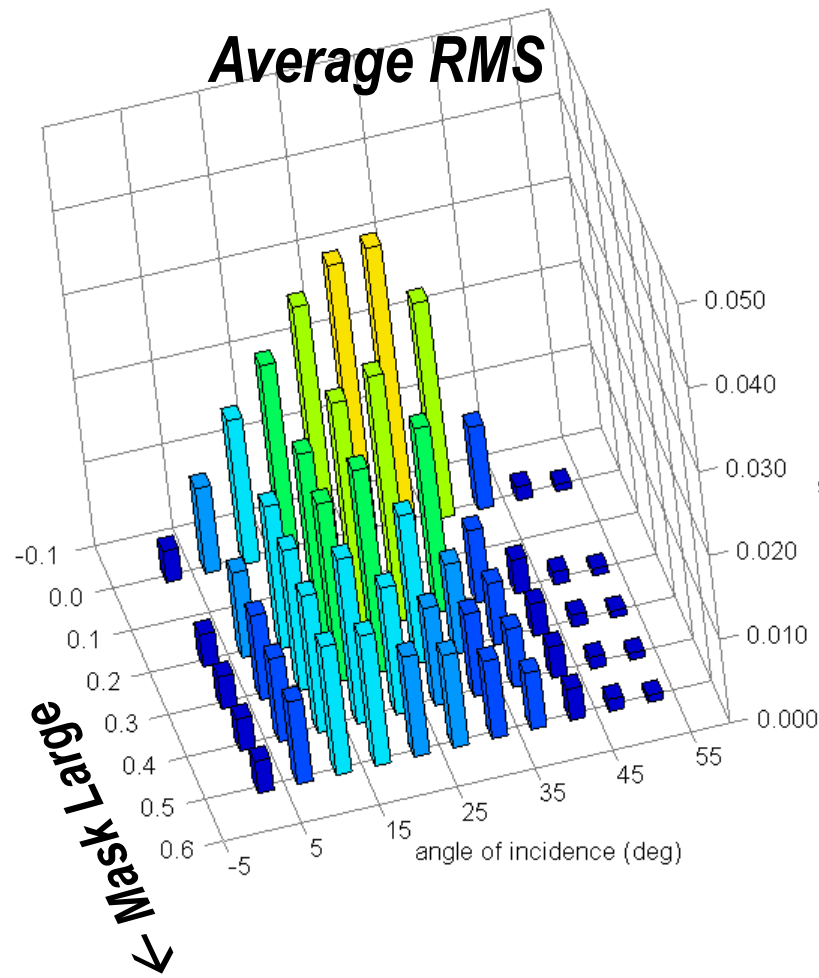
# Flat mask cases: Intensity

**Mask Case = 1.00**

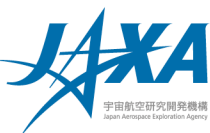
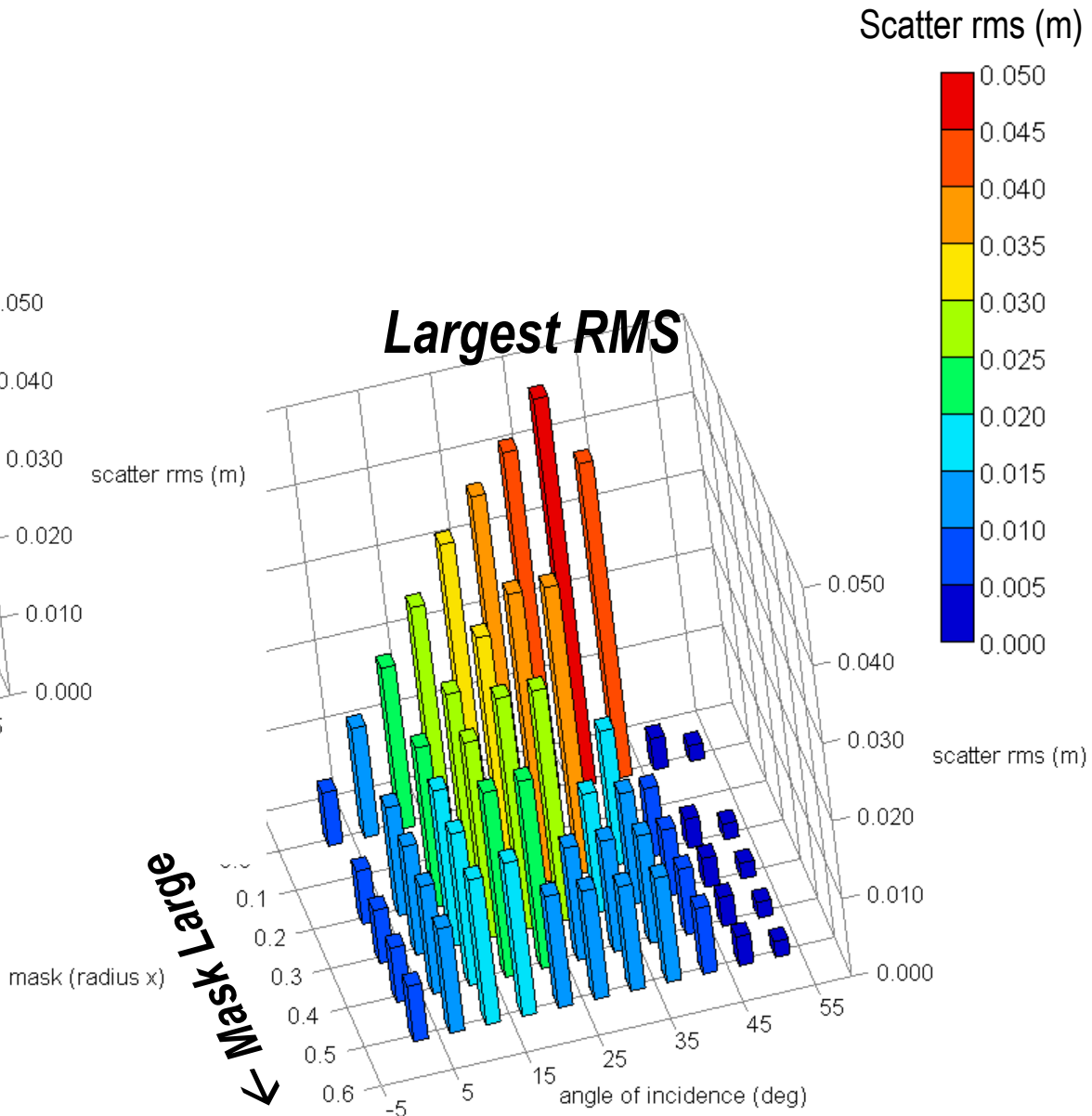


# Delta mask cases: Scatter RMS

### Average RMS



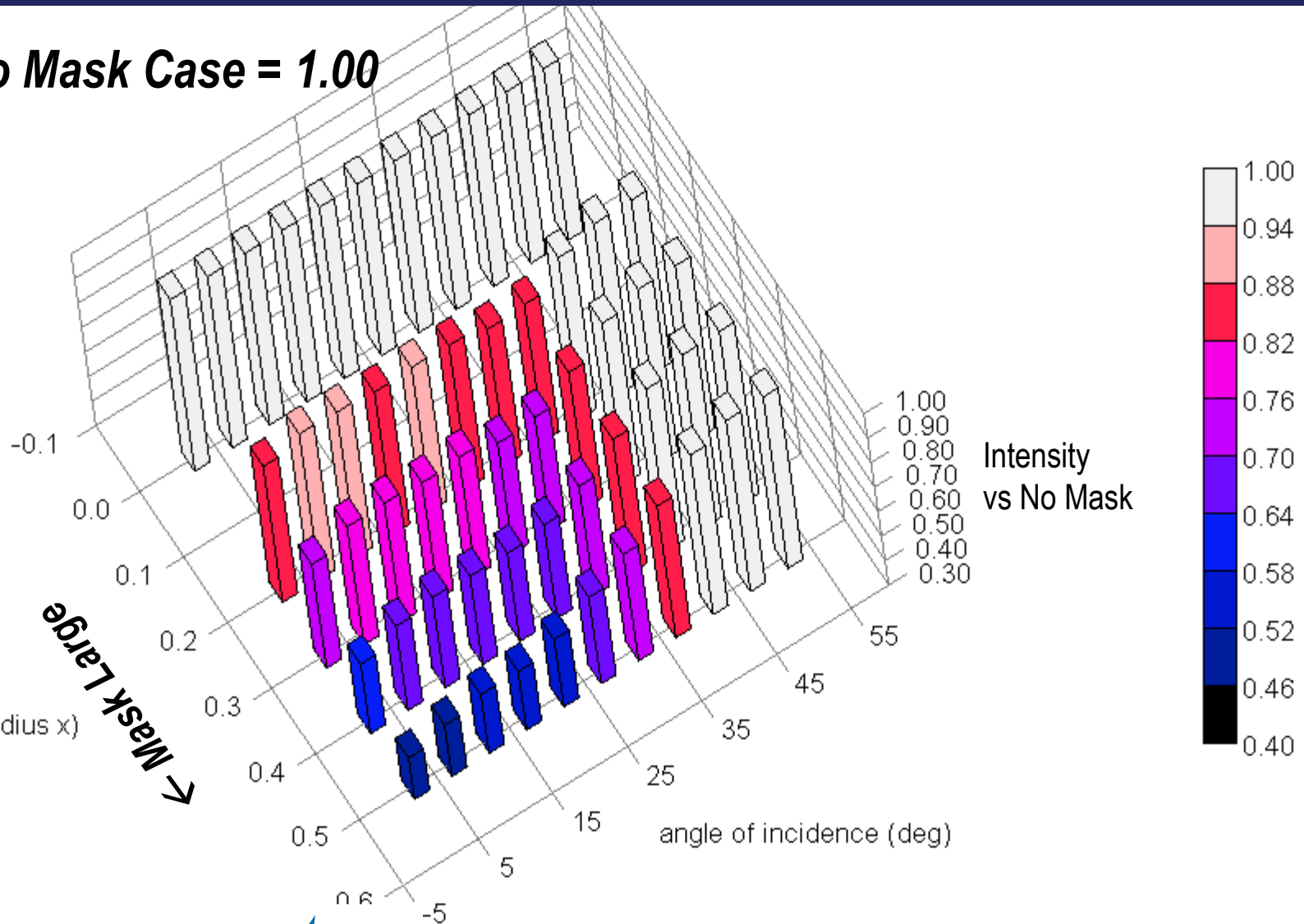
### Largest RMS



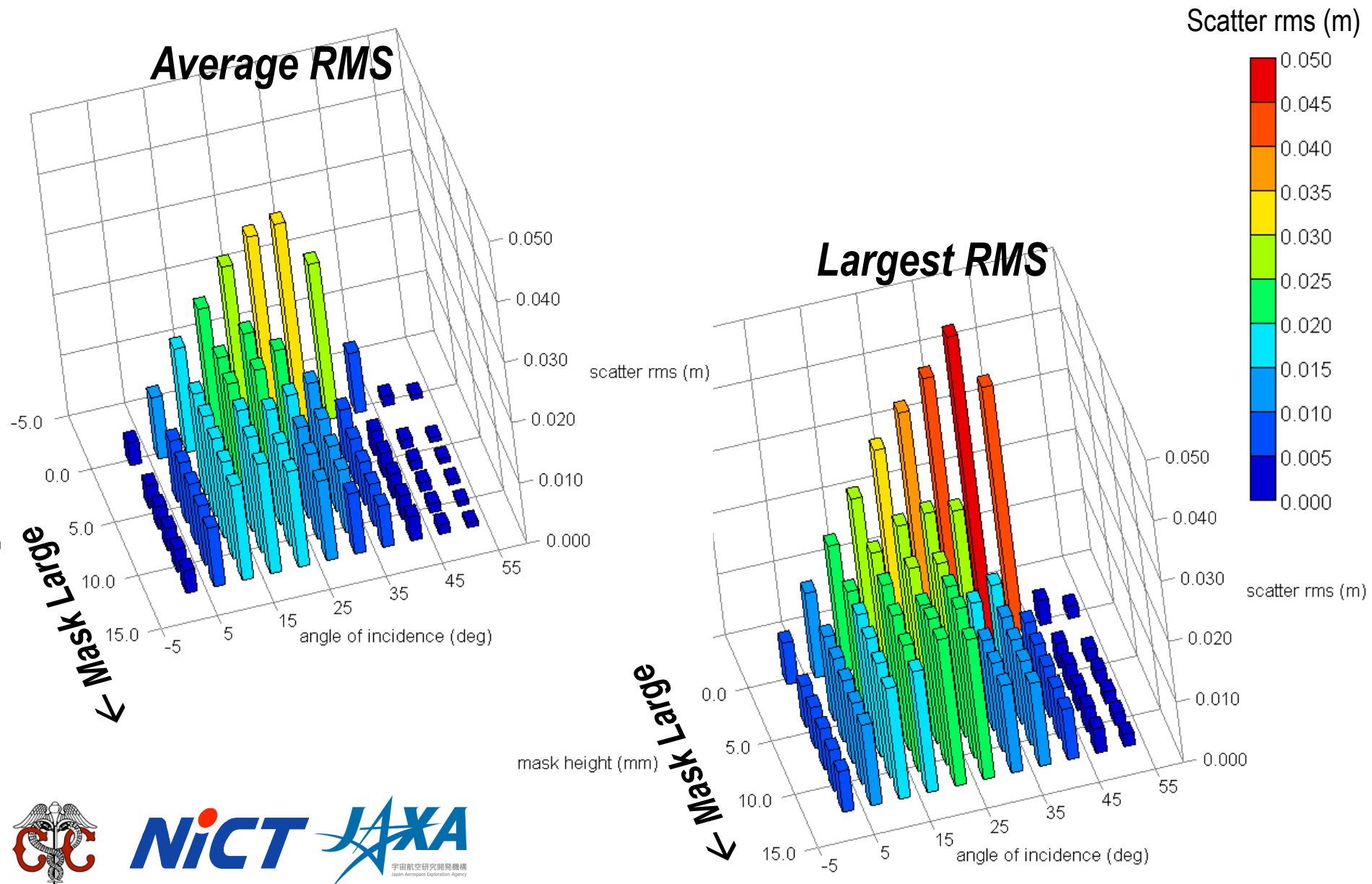
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# Delta mask cases: Intensity

**No Mask Case = 1.00**



# Fin mask cases: Scatter RMS



# Fin mask cases: Intensity

**No Mask Case = 1.00**

