



# Optical Laser time transfer and high repetition rate monostatic SLR

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## Increase Repetition Rate

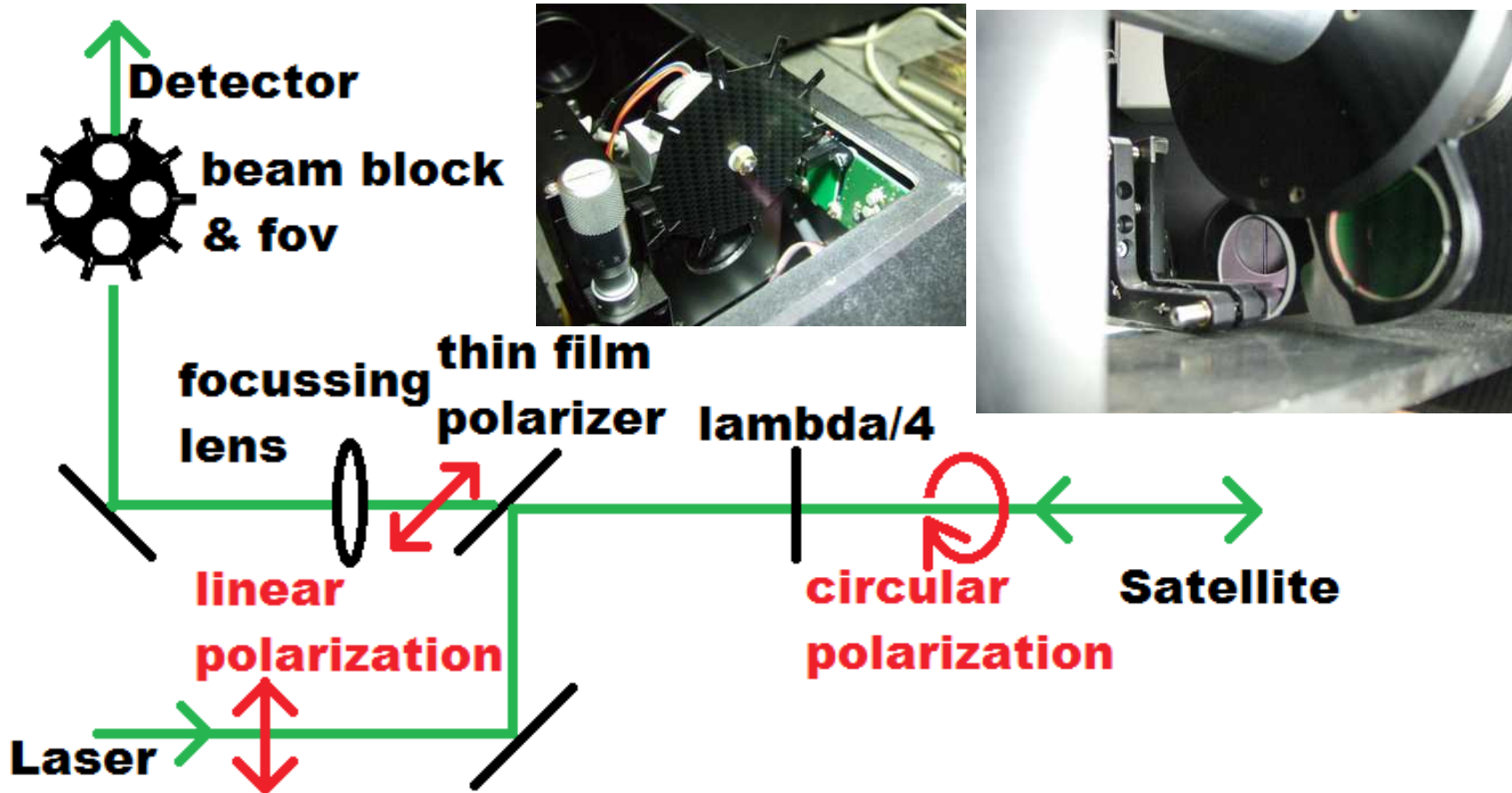
- Return Rate  $> 0.01$  to allow daytime automatic SLR
- Precision  $< 1$  mm for all ILRS targets, but reduce systematics  $\rightarrow$  going optical?

## NIR ranging (already shown @ 2013 ILRS WS)

- Eye-Safe in combination with telescope aperture
- Improved link efficiency

## Time Transfer capability

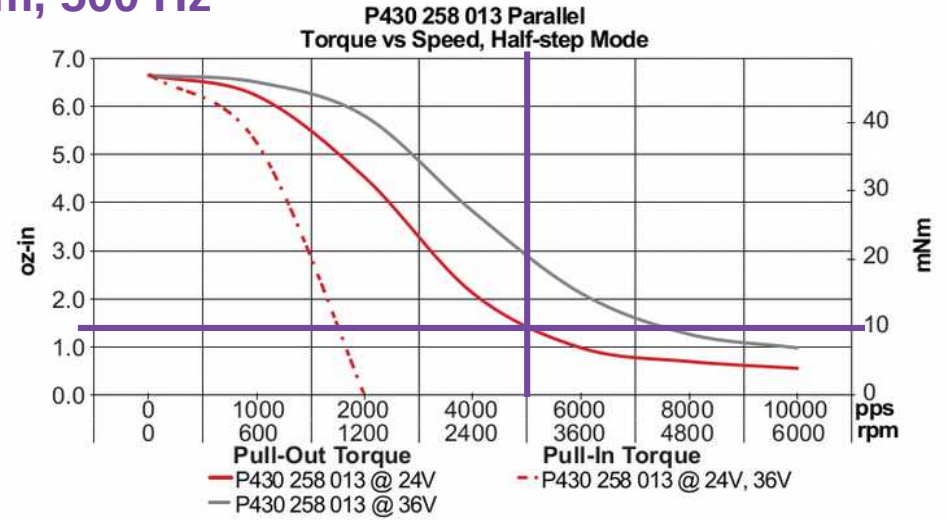
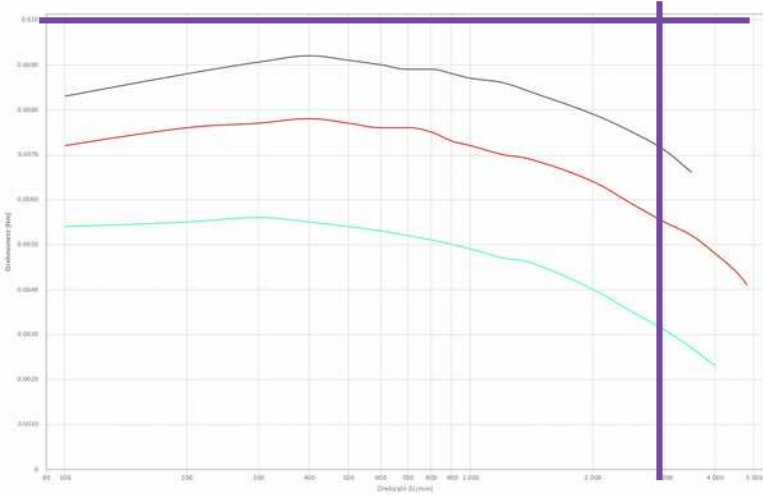
- Laser Fire synchronization to local Timescale
- Improved system stability by new Start & Event-Timer electronics (CTU Prague)



**Signal is forced to be circular polarized!**

# T/R switch - stepper motor -

10 mNm, 500 Hz



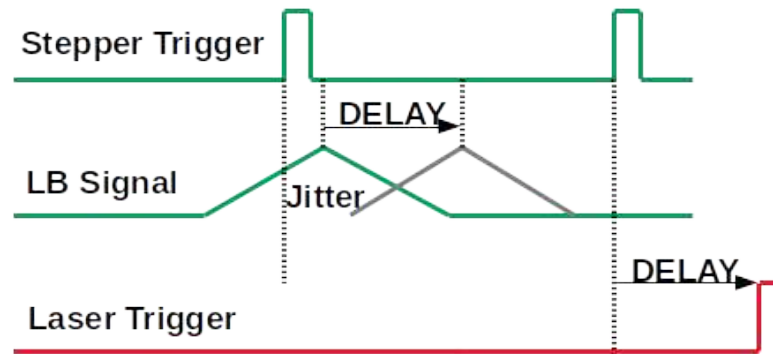
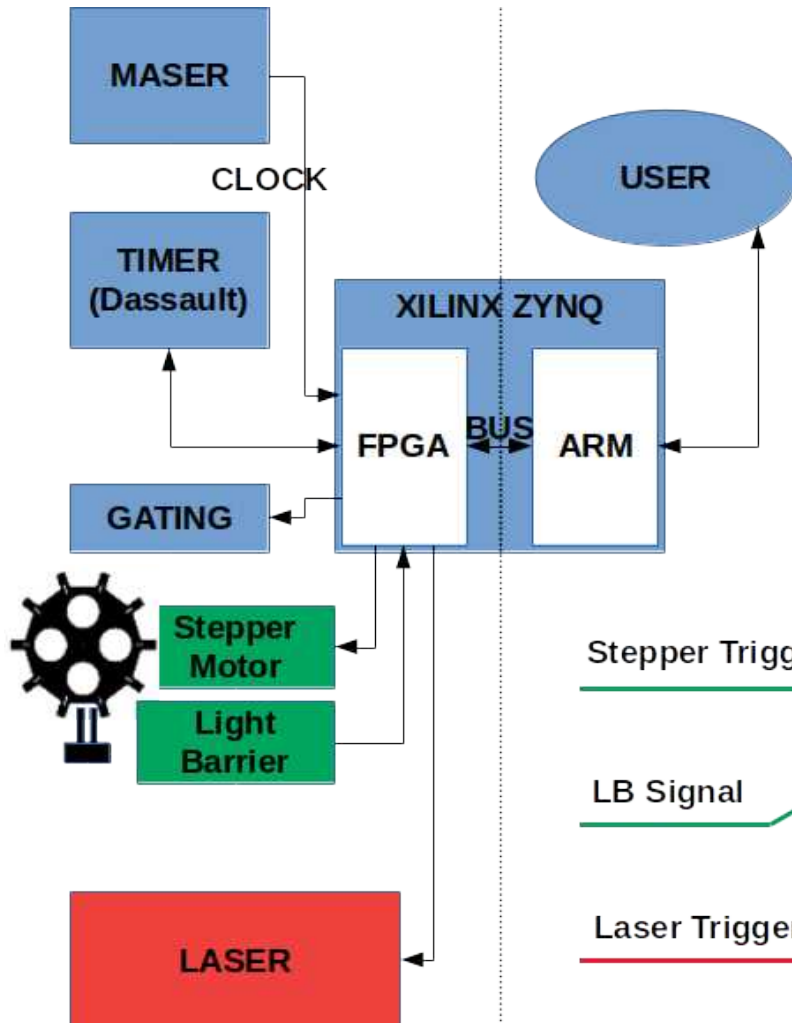
now



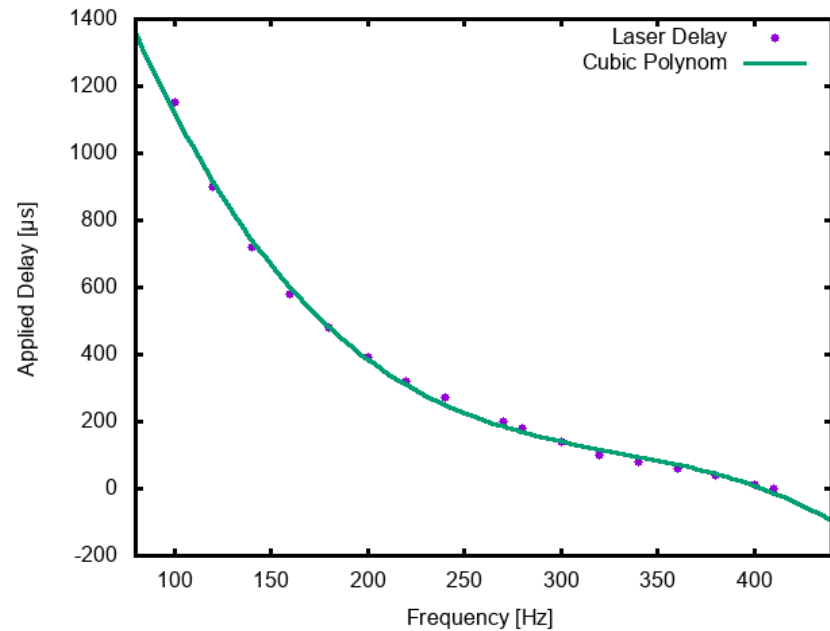
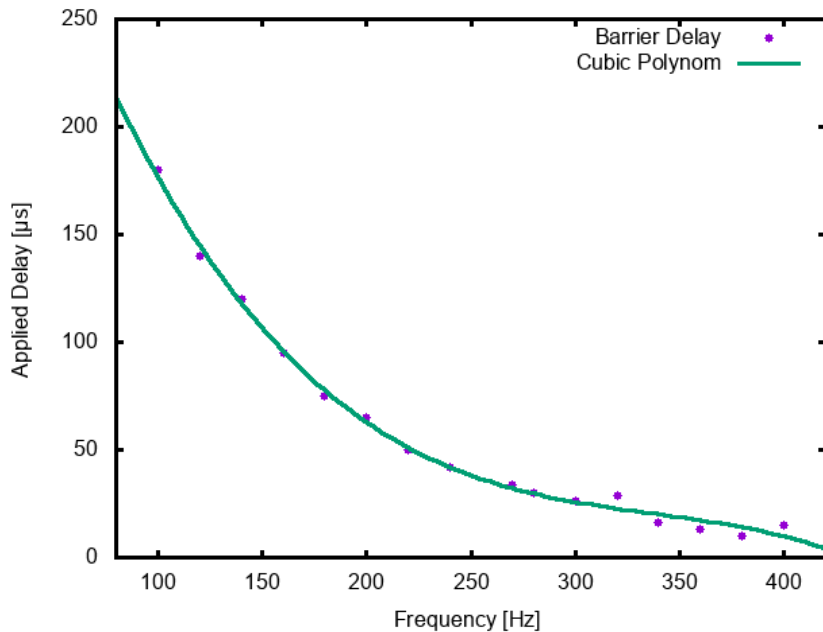
future



# T/R switch - electronics & timing -

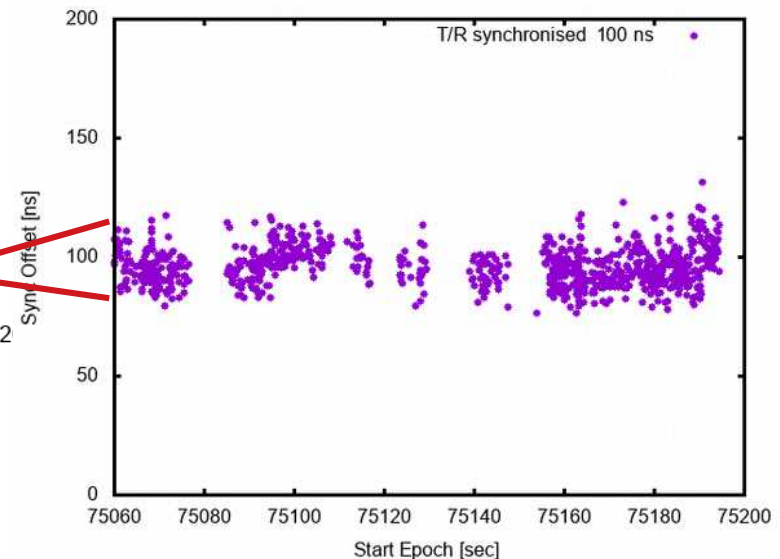
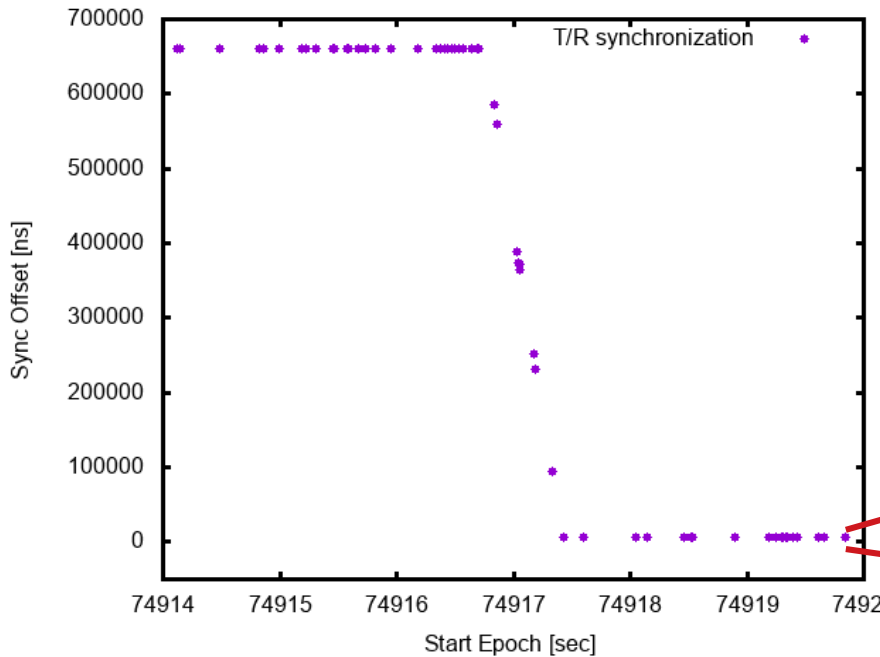


- Light Barrier & Laser Delay frequency dependent
- Collision avoidance requires frequency shift  
→ adjusted using cubic fit



- Further Requirement: Passive Mode-Locked Oscillator lasing continuously → AOM implemented

## Synchronization of Laser Fire Times to external Time-Base via phase shift of stepper trigger frequency



- Slew rate  $\sim 1$  ms/sec
- P-P Jitter  $\sim 25$  ns



**Max. Laser pulse energy:** 1.2 mJ (used: ~0,7 mJ)  
**Repetition Rate:** 400 Hz  
**Ranging Wavelength:** 1064 nm  
**Telescope aperture:** 0.75 m (transmit & receive)

**Data clipping:**

2.0 for internal calibration (spurious tail)  
2.2 for external reference calibration (Gaussian)  
2.2 for satellites

**Data rejection:**

> 2.5 mm precision  
< 10 FR points per NP  
> 10% Return Rate



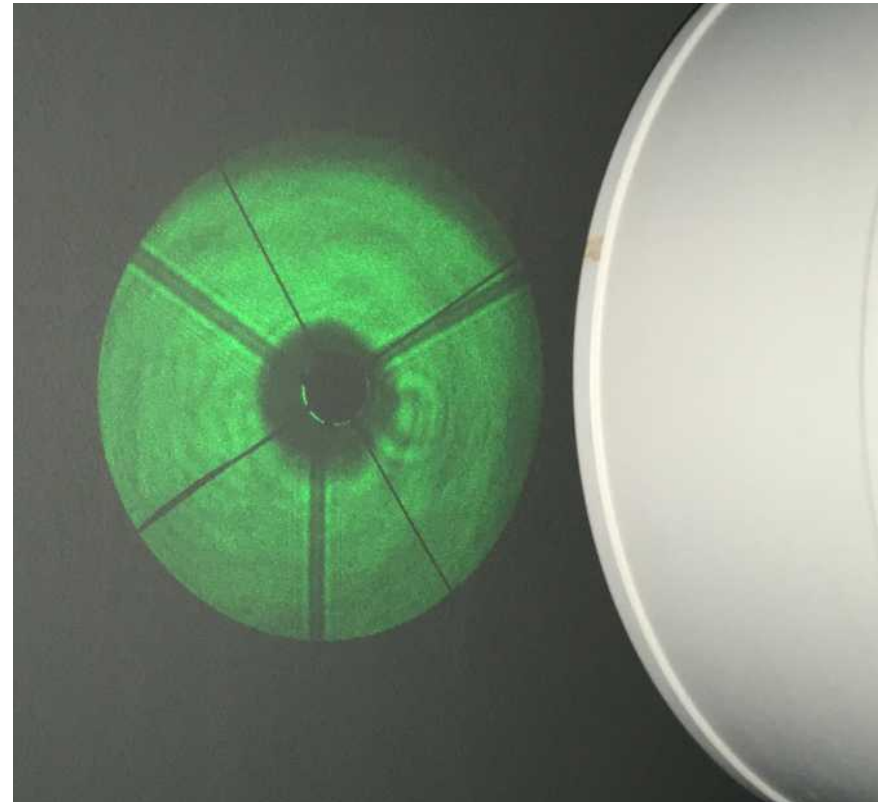
- Measure laser pulse energy of brightest spot of beam profile @ telescope output:  
532 nm →  $1.5e-8$  J/cm<sup>2</sup>  
1064 nm →  $0.5e-7$  J/cm<sup>2</sup>

- *Max. ANSI Z136.1 (2014):*  
532 nm →  $1e-7$  J/cm<sup>2</sup>  
 $1064$  nm →  $2e-6$  J/cm<sup>2</sup>

- *Max. IEC 60825-1 (2014):*  
532 nm →  $1.9e-9$  J/cm<sup>2</sup>  
 $1064$  nm →  $2.5e-8$  J/cm<sup>2</sup>

- **Factor 2 only!**

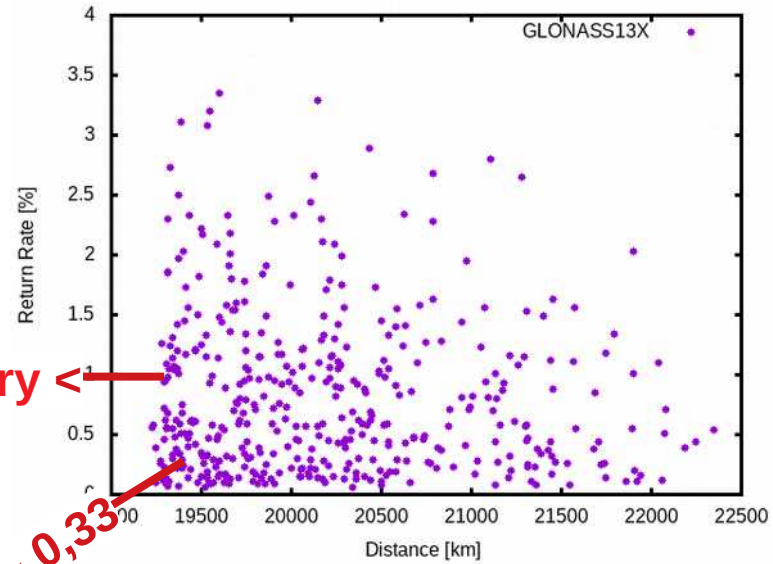
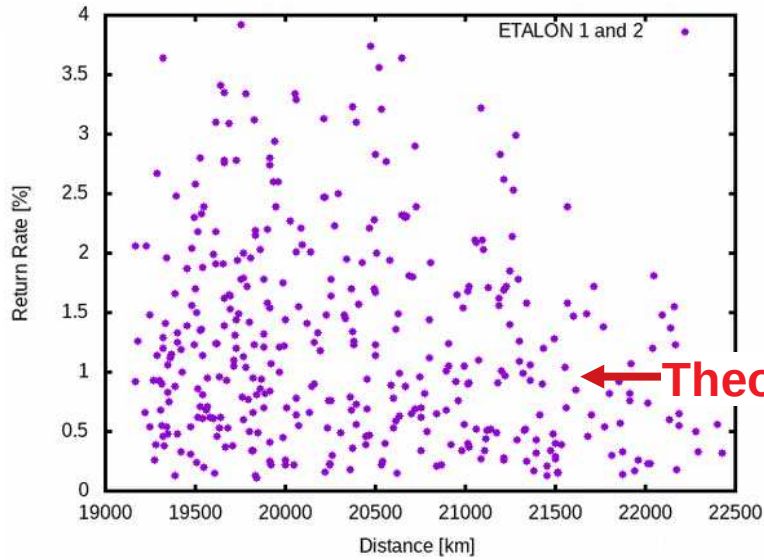
**NO MPE! - CLASS 1:** Starring 10 sec with  
laser beam



into the

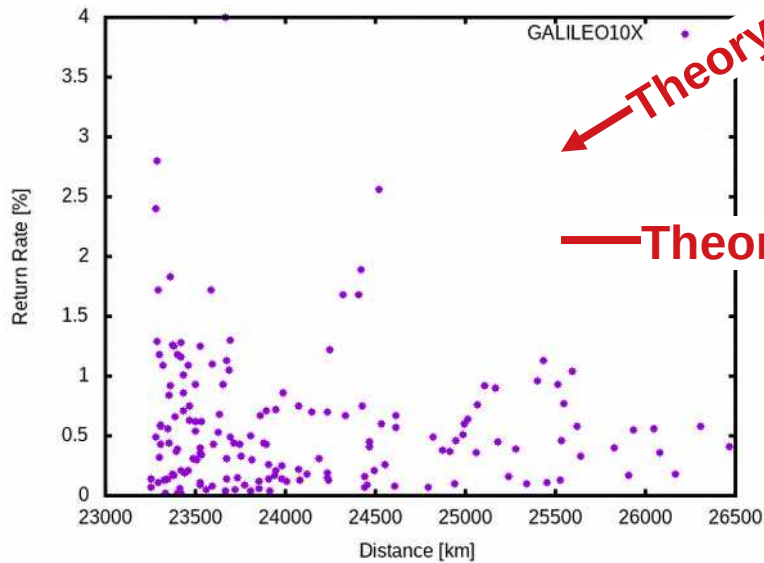


# High Rep-Rate Ranging - heo performance -

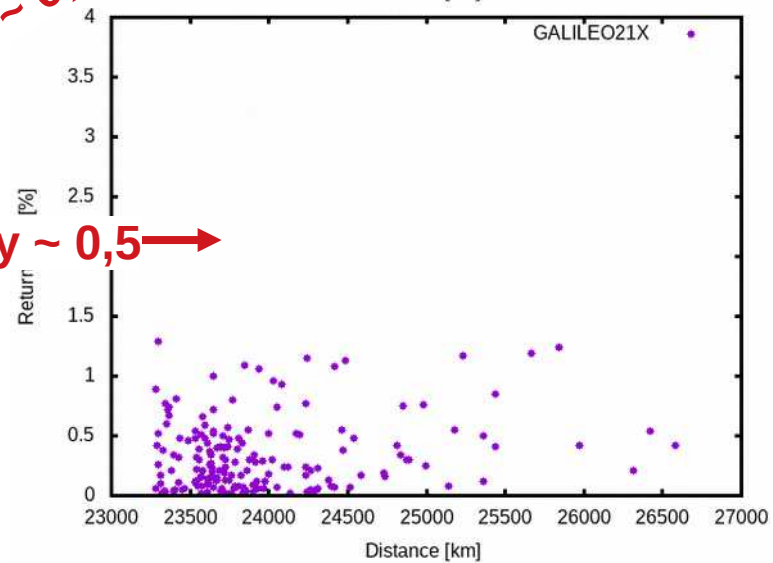


← Theory ←

← Theory ~ 0,33

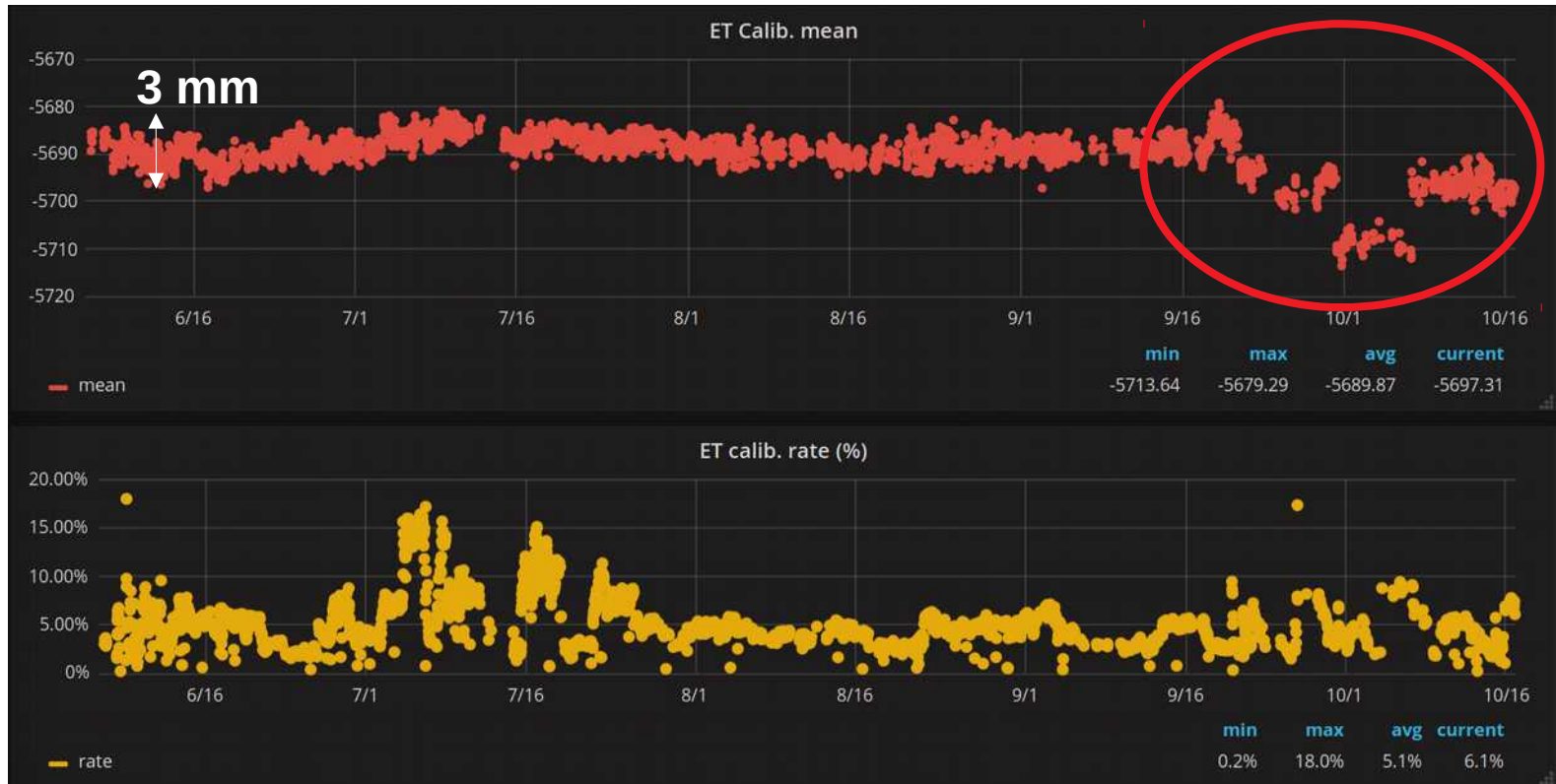


→ Theory ~ 0,5 →





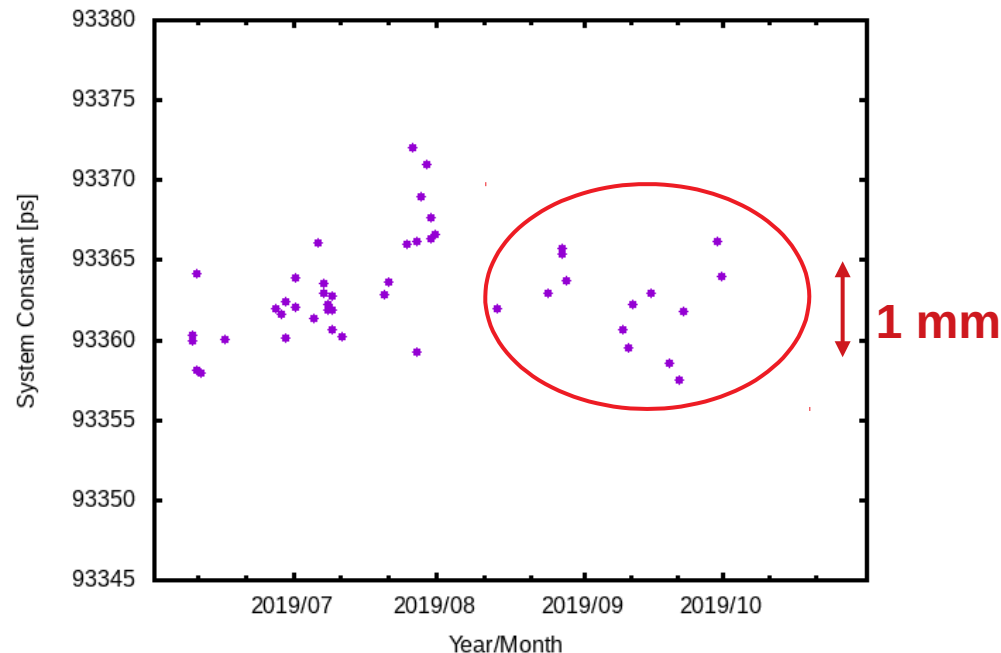
# Calibration - absolute delay stabilization -



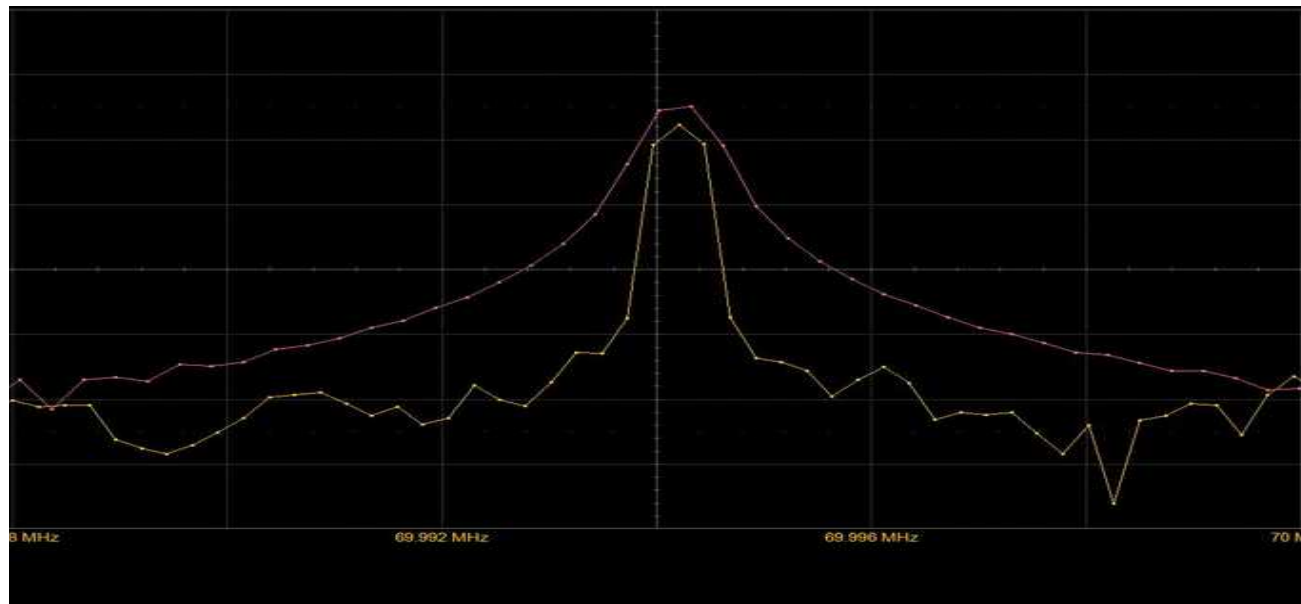
P-P Variation of internal calibration 15 ps (<3mm) over month!

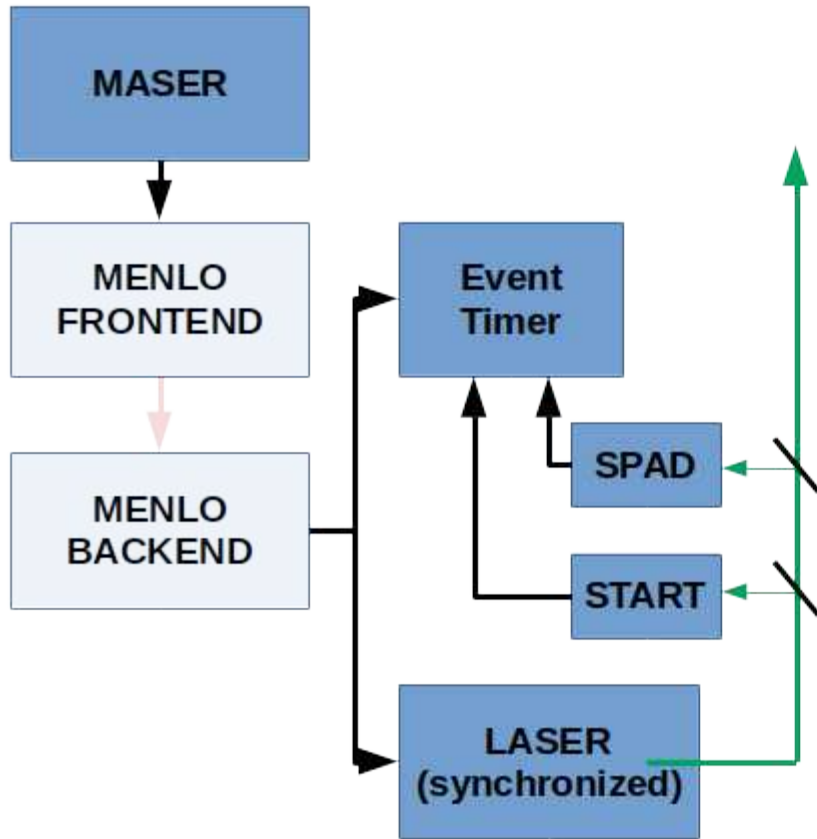


- Use telescope internal calibration to interpolate external target reference measurements
- Advantage:
  - No telescope positioning required
  - Determination of system constant possible

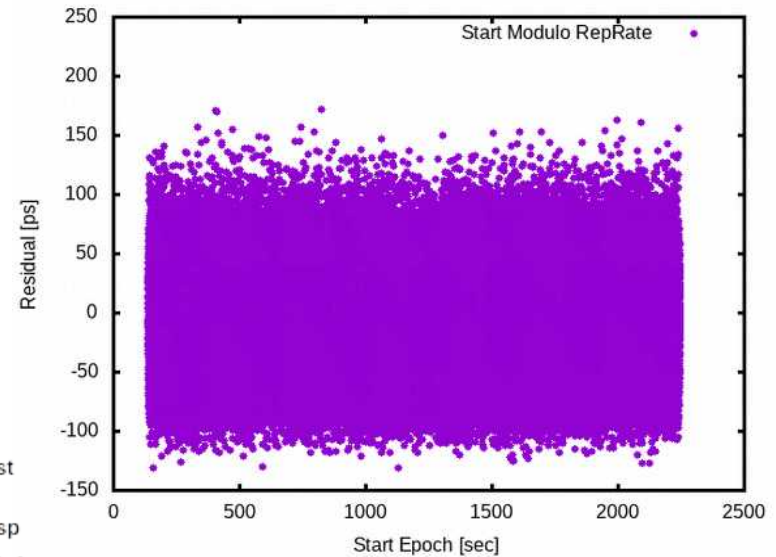


- Passive mode-locked laser oscillators provide high spectral purity
- Lock repetition rate to reference frequency to stabilize start epoch or verify timing electronics





START EPOCH	STOP EPOCH	RANGE (Calibration)
.	.	.
.	.	.
.	.	.
.	.	.
.	.	.



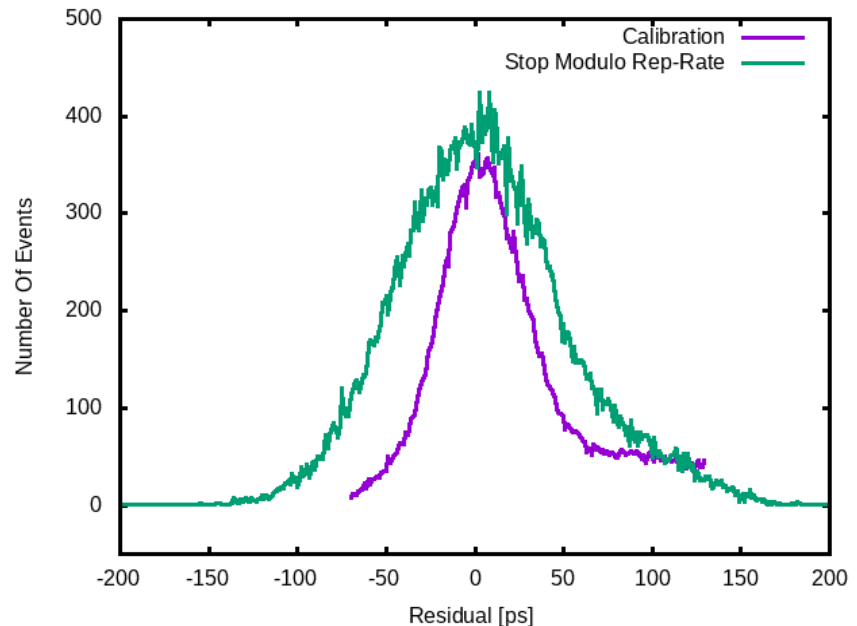
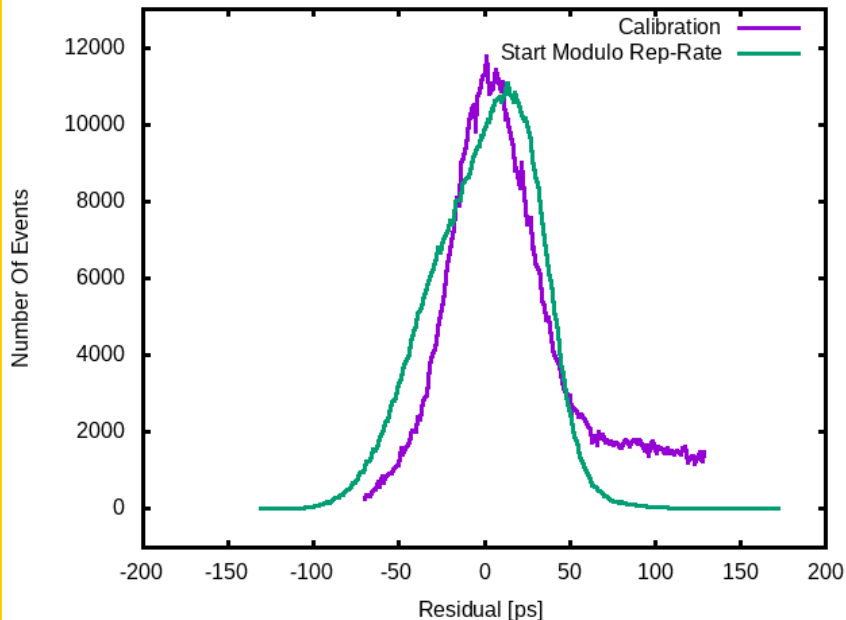
$$\text{START} = n * (\text{RepRate})^{-1} + \Phi_{\text{Start}} + d\Phi_{\text{common}} + d\Phi_{\text{err\_st}}$$

$$\text{STOP} = n * (\text{RepRate})^{-1} + \Phi_{\text{stop}} + d\Phi_{\text{Common}} + d\Phi_{\text{err\_sp}}$$

$$\text{RANGE} = \text{STOP} - \text{START} = \Phi_{\text{stop}} - \Phi_{\text{Start}} + d\Phi_{\text{err\_sp}} - d\Phi_{\text{err\_st}}$$



- Identify errors common to both channels
- Independent validation of Start & Stop channel
- Preliminary result!



- Future: start epoch generation @ “sub”-ps accuracy with two-way link to maser



- Galileo2XX tracking: Is 532 nm front coating problematic for 1064 or 850, is there a coating?
- Do other stations have similar experience with HEO ranging?
- What is general opinion about 1064 nm ranging?
- Do other stations apply data rejection criterion, which one?
- Sentinel ranging at 1064 nm?
- Should we sent passage or passage segment data?